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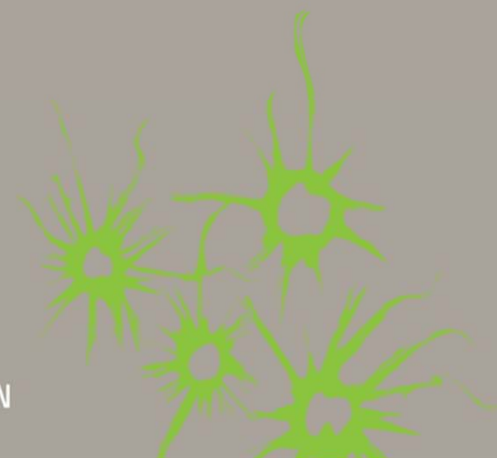
INFRARED SPECTROSCOPY ON MINERALS AND ROCKS

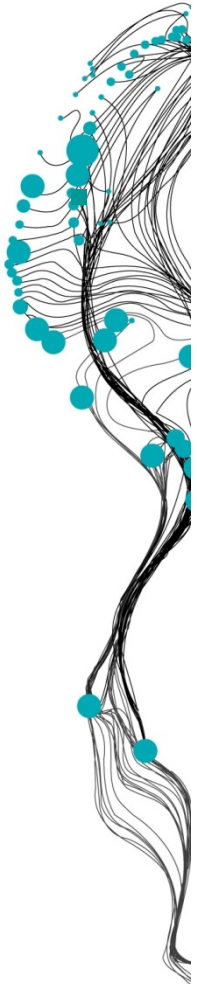
CHRIS HECKER (ITC)

JULY 2013



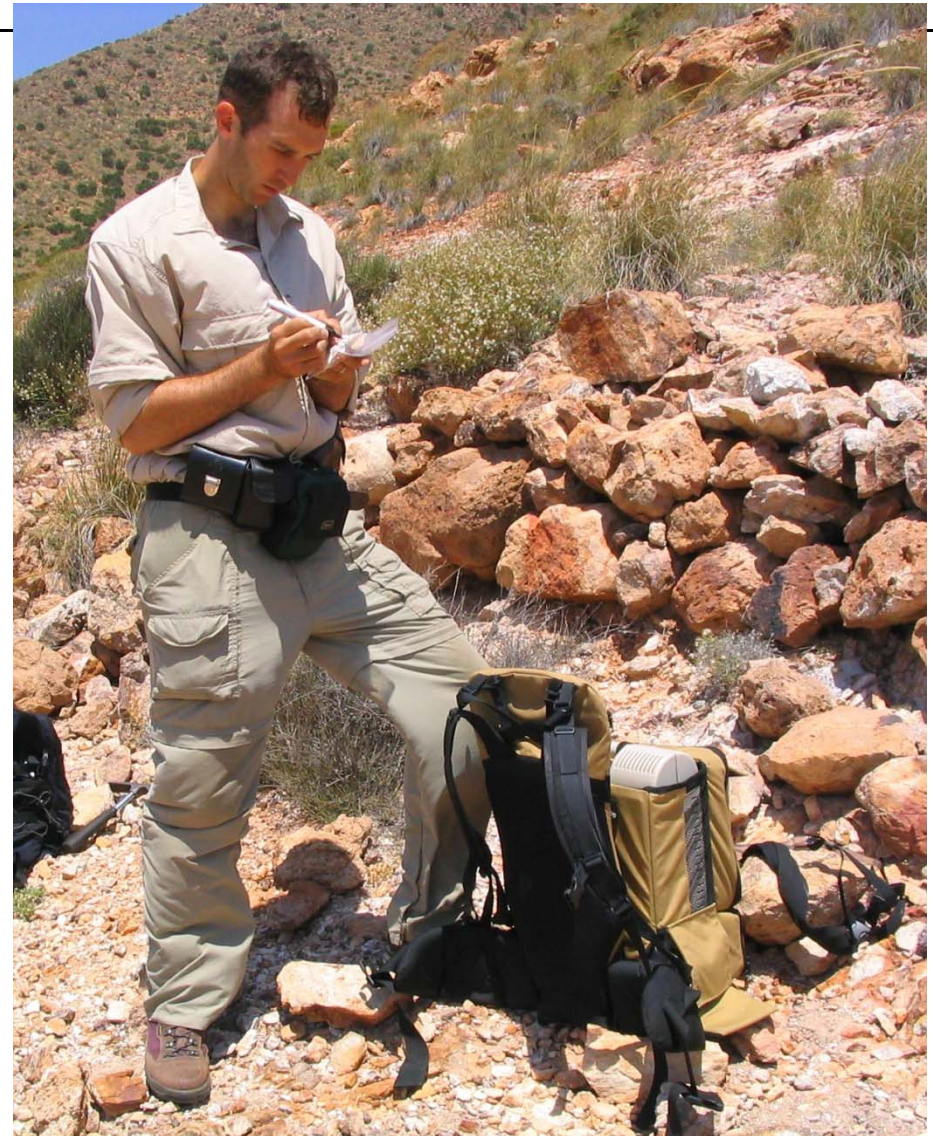
FACULTY OF GEO-INFORMATION SCIENCE AND EARTH OBSERVATION





INTRO: CHRIS HECKER

- CV
 - B.Sc. Earth Sciences (1996)
 - M.Sc. Earth Sciences (1999)
 - University of Basel, Switzerland
 - PhD Thermal RS (2012) ITC
- Lecturer Geologic Remote Sensing (70%)
- Researcher Geologic Remote Sensing (30%)
- Areas of interest:
 - Geologic RS
 - Thermal RS
 - Imaging Spectrometry
- Email: c.a.hecker@utwente.nl





INTRO UT-ITC (IN A NUTSHELL)

- Mission:
development and transfer of knowledge in
geo-information science and earth
observation
- Academic level:
PhD/MSc/Master/Postgraduate Diploma
- Target group:
young and mid-career professionals, and
scientists from developing and emerging
countries, increasingly professionals
from industrialised countries
- Framework
international development cooperation



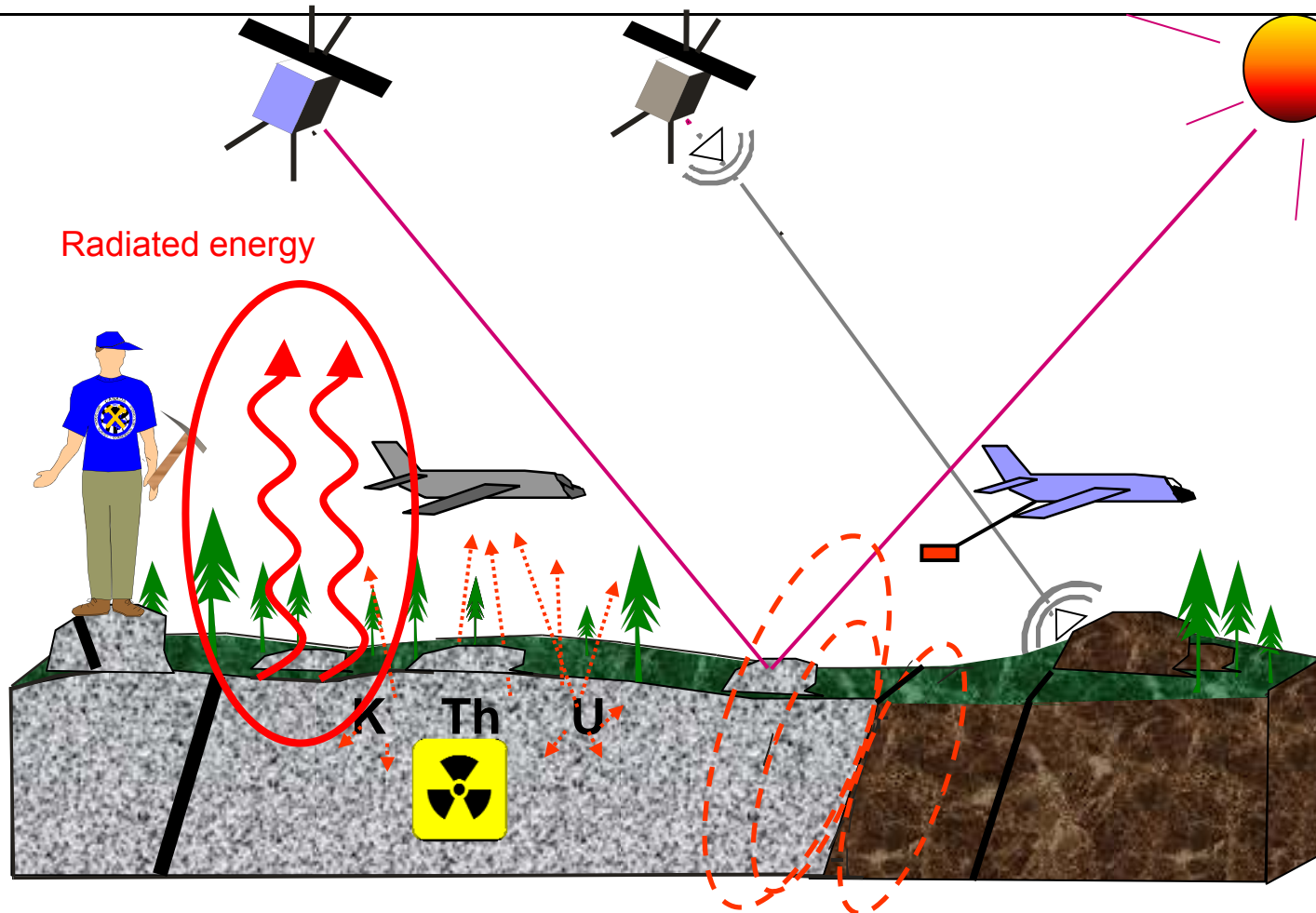


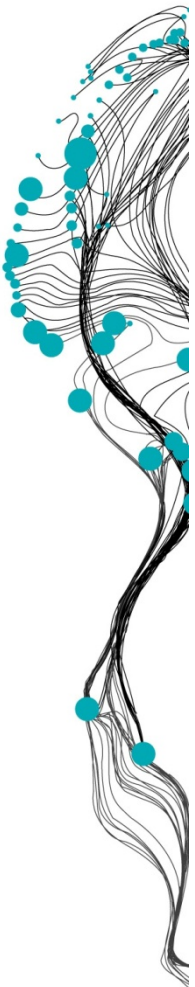
OUTLINE OF LECTURE

- ***SWIR vs TIR***
- Emissivity spectra of Minerals and Rocks
- Ground-based setups
 - Laboratory
 - Field
- Mapping methods
- Case study: TIR+PLSR

THERMAL INFRARED

ONE OF THE TOOLS IN THE MODERN EARTH SCIENTIST'S TOOLBOX

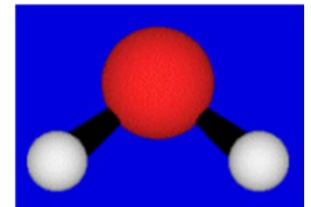




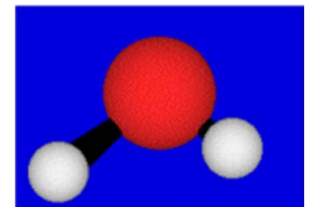
FUNDAMENTAL VIBRATIONAL FREQUENCIES

- The bonds in a molecule or crystal lattice are like springs with attached weights: the whole system can vibrate
- Different types of vibration possible
- Each have different Energy levels
- Combination of absorption features can be diagnostic
- Examples:
 - Al-OH, 2.20 μm
 - Mg-OH, 2.3 μm
 - Ca-CO₃, 2.32-2.35 μm
 - Si-O, ~ 9-10 μm

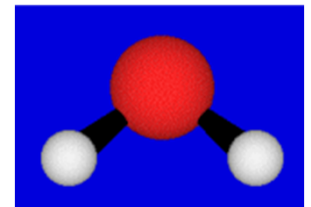
Stretching

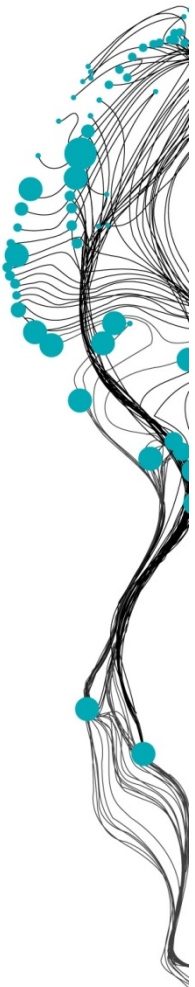


Asymmetric Stretching



Bending





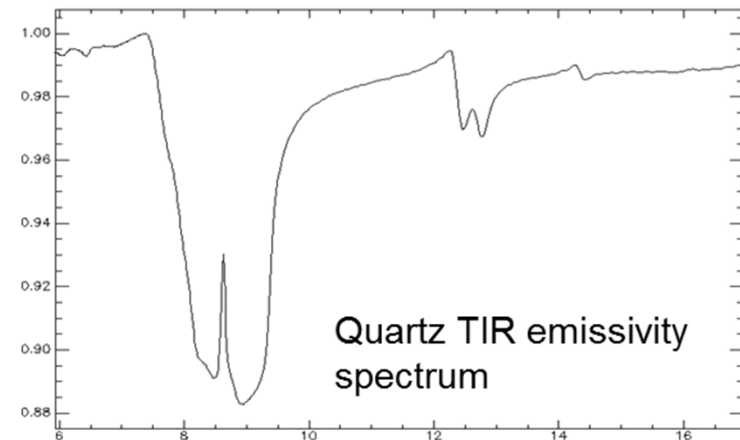
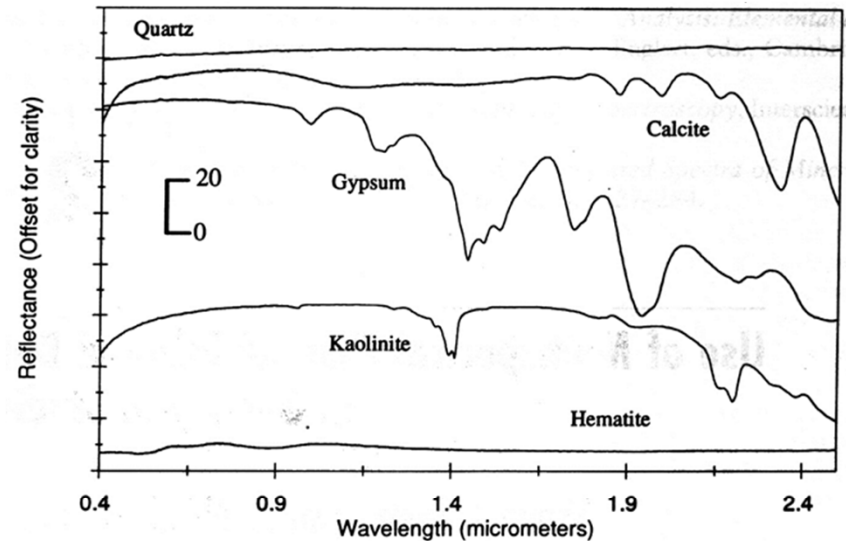
WHY EMISSIVITY SPECTRA?

- VIS-SWIR features in minerals with:
 - Iron
 - Hydroxyl / water
 - Sulfates
 - Carbonates
 - Phosphates
- No SWIR features for non OH-bearing Silicates => TIR emissivity spectra needed

Rule of Thumb:

- VNIR/SWIR shows alteration products
- TIR shows differences in rock composition

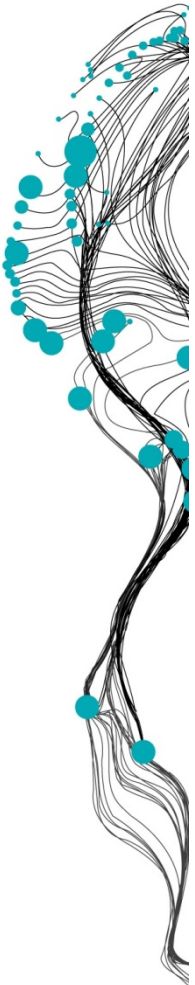
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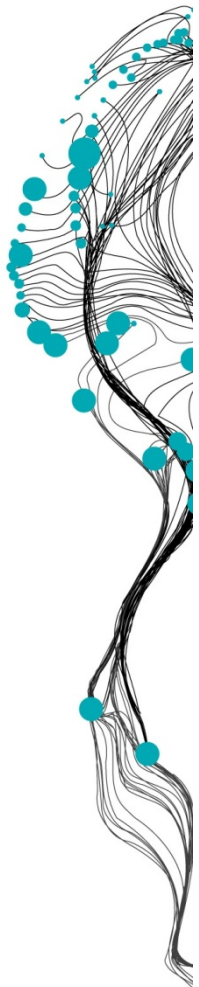
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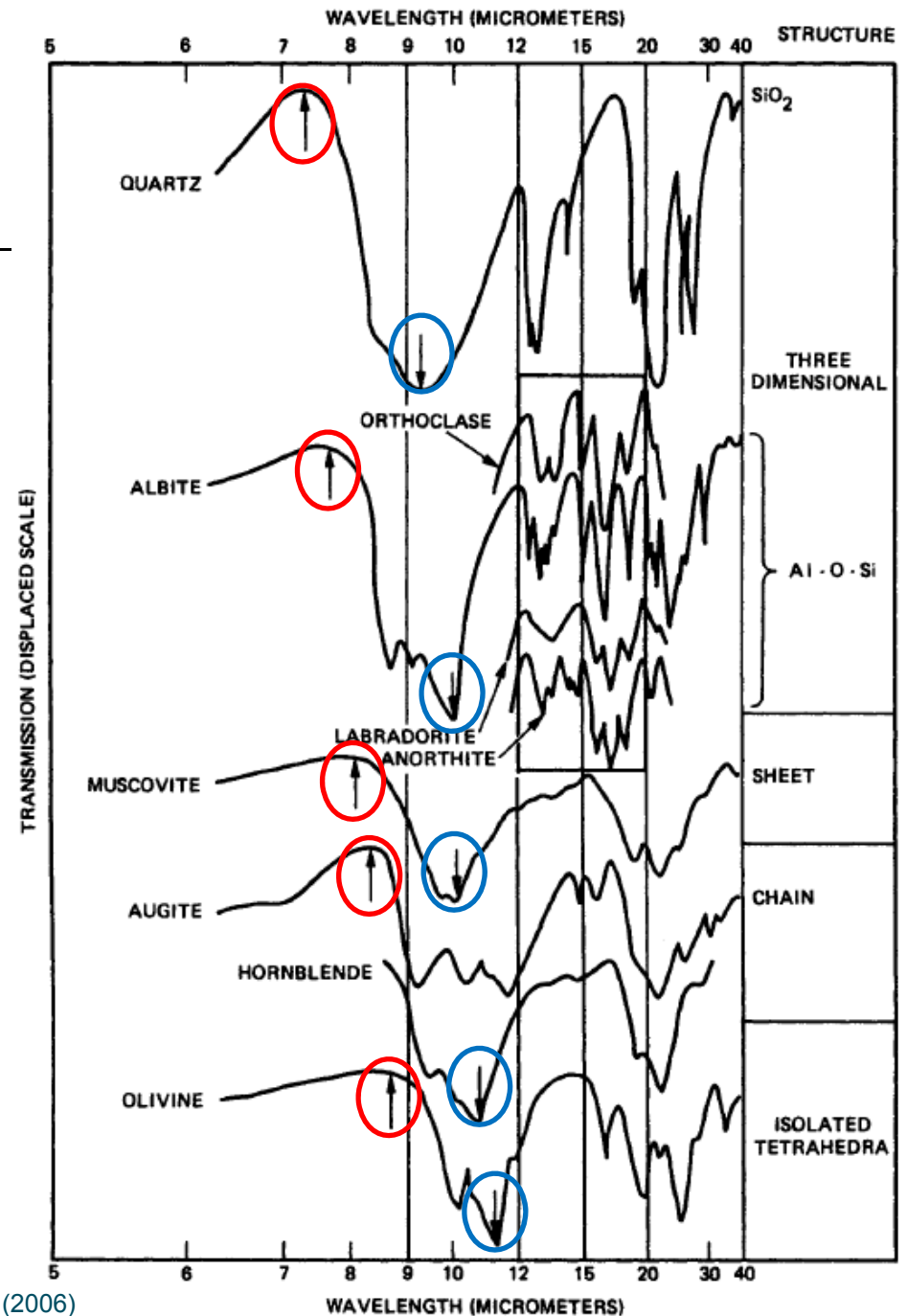
TYPICAL MINERAL SPECTRA

- Reststrahlen feature
 - Strong reflection peak / emission minimum due to fast change in refractive index and center of strong absorption band.
 - Causes emissivity low
 - Shape is diagnostic for silicates and other minerals
- Christiansen frequency
 - Wavelengths where refractive index is close to unity => little scattering
 - If not in absorption band, causes high transmission and low reflectance
 - Visible as emissivity maxima in spectra



TYPICAL MINERAL SPECTRA (CONT'D)

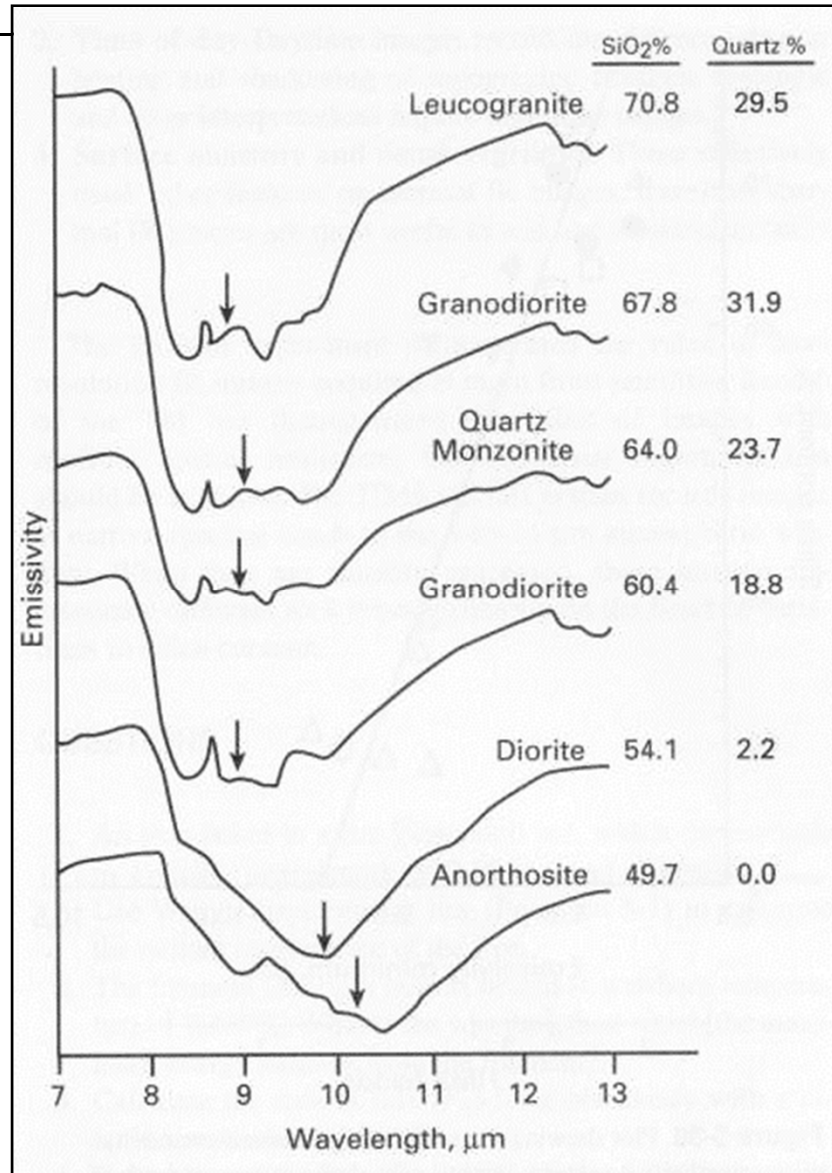
- Mineral (transmission) spectra showing:
 - Christiansen features (**up arrows**)
 - Reststrahlen features (**down arrows**)
 - Positions shift to longer wavelengths with decreasing Si-O₄ tetrahedra polymerization.





TYPICAL ROCK SPECTRA

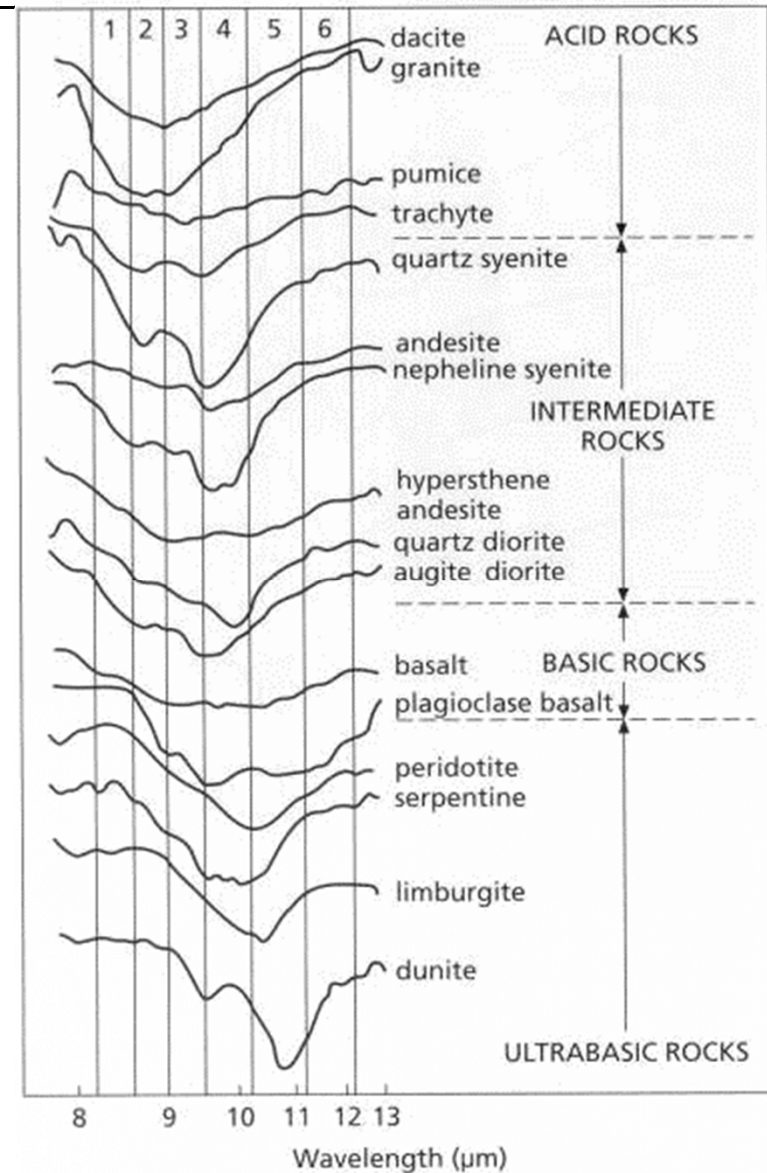
- Rock spectra usually more complex than mineral spectra
- Rock spectra combine features of their main mineralogy
- Acidic rocks show reststrahlenband at lower wavelength than basic rocks
- Change in emissivity minimum can be used for mapping igneous rocks of variable SiO₂ content

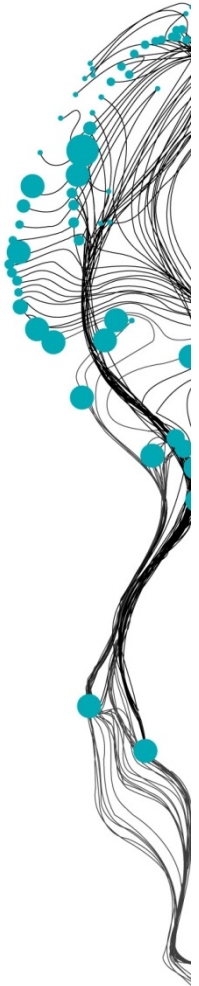




TYPICAL ROCK SPECTRA (CONT'D)

- Multi-band thermal systems can help distinguish different rock types and compositions
- Vertical lines and numbers indicate 6 bands of the Thermal Infrared Multispectral Scanner (TIMS)





TYPICAL ROCK SPECTRA (CONT'D)

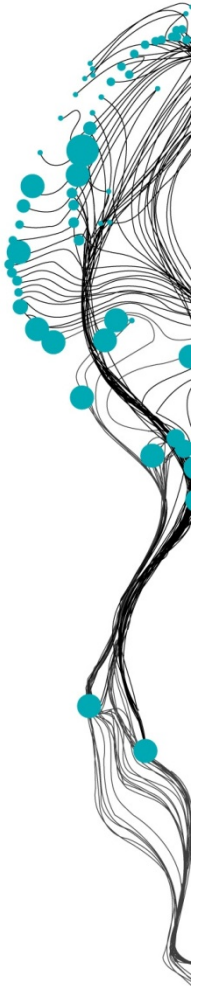
What can we do with it in rock / soil mapping?

- **Christiansen frequency**

- Exact position not diagnostic in mixtures
- Generally high emissivity around 7.5 μm (and 12 μm) useful in T&S.

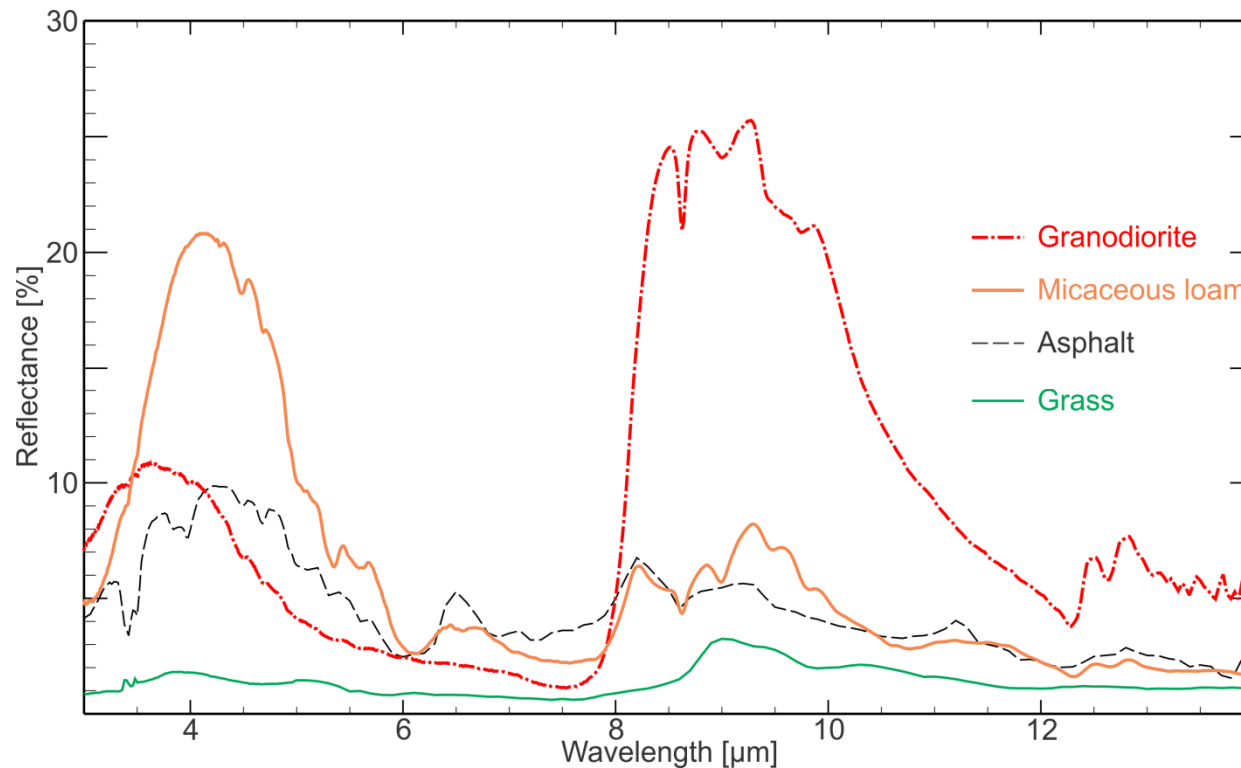
- **Reststrahlen feature**

- General position / shape can give hint in multispectral mapping (e.g., silica%).
- “Deciphering” of reststrahlen feature used for quantitative analysis in spectroscopy (e.g., PLSR or unmix)
- Reststrahlen feature of rocks are great for practicing field spectroscopy (before attempting e.g., plants)



TYPICAL ROCK SPECTRA (CONT'D)

- Spectral contrast of rocks much higher than in soils or vegetation
- Example of DHR spectra from ASTER speclib





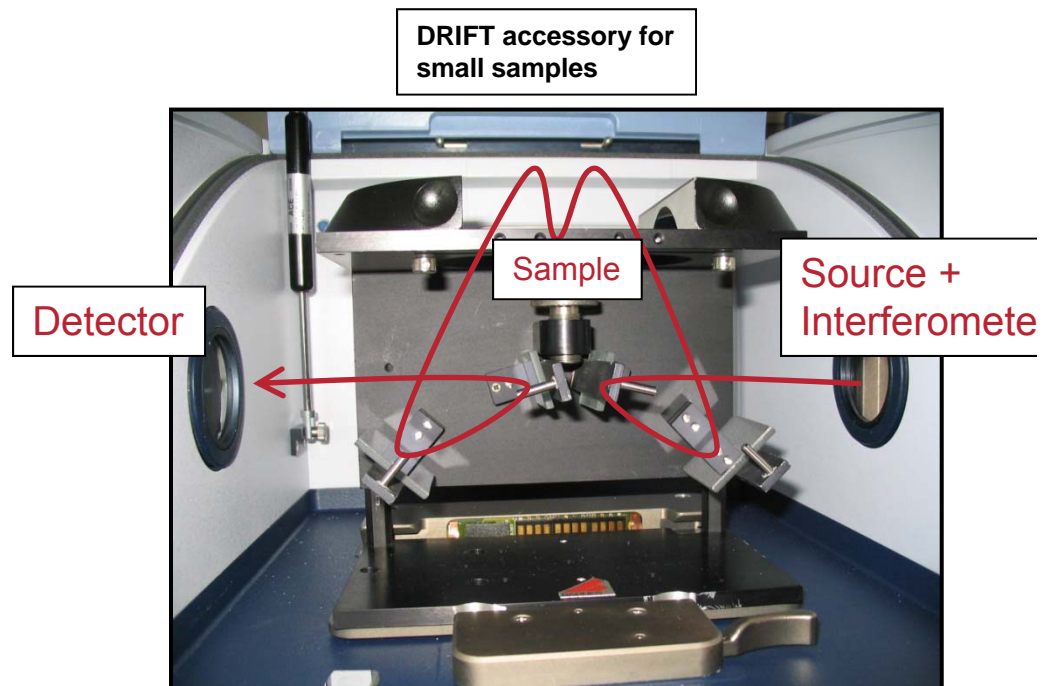
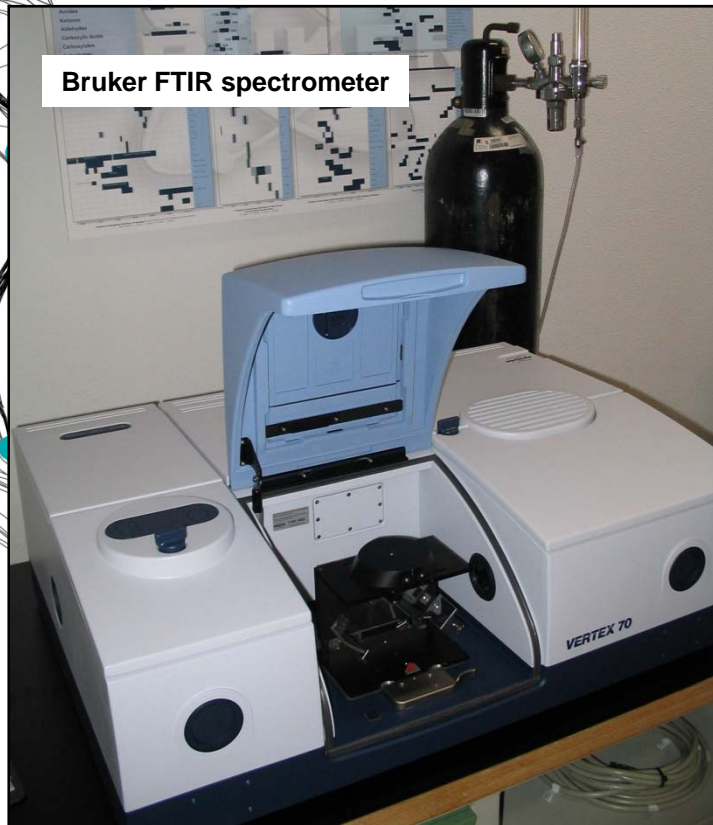
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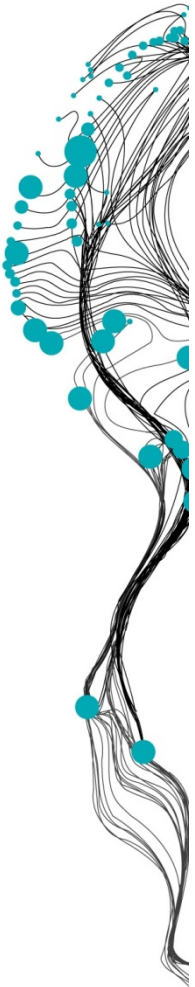
THE SPEC LAB FAMILY PORTRAIT



LABORATORY FTIR BRUKER & DRIFT

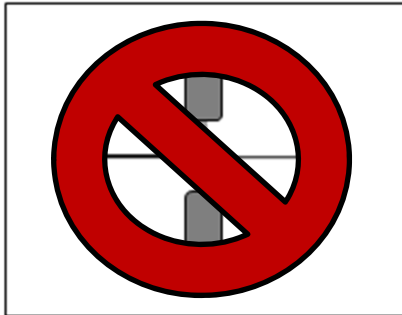


LABORATORY – SAMPLE CONSIDERATIONS

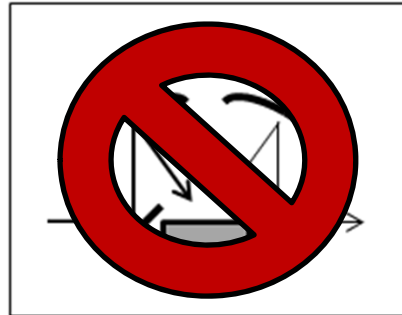


- Designed for small powder samples
- Sampling spot and space too small for most geologic samples

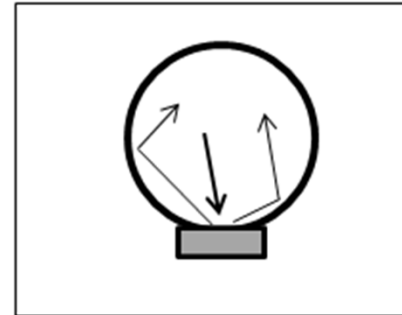
LABORATORY – GEOMETRY CONSIDERATIONS



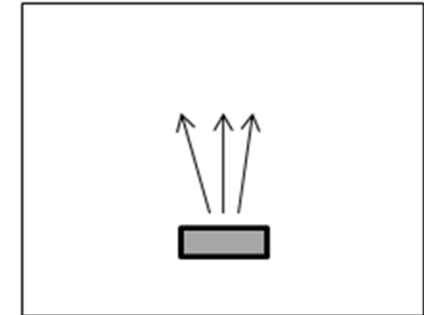
Transmission



Bi-dir refl



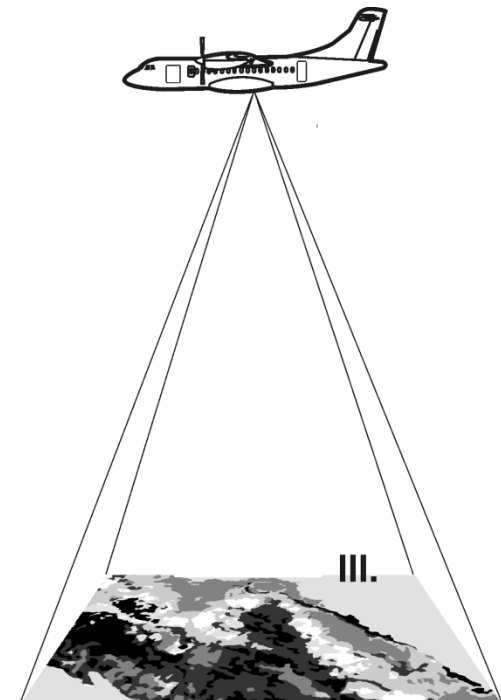
Dir-hem refl



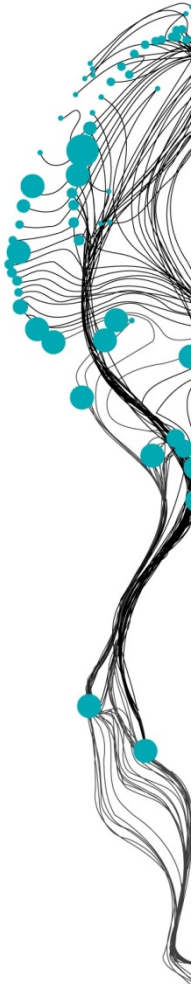
Emission

Quantitative comparison with RS data: DHR or Emission only

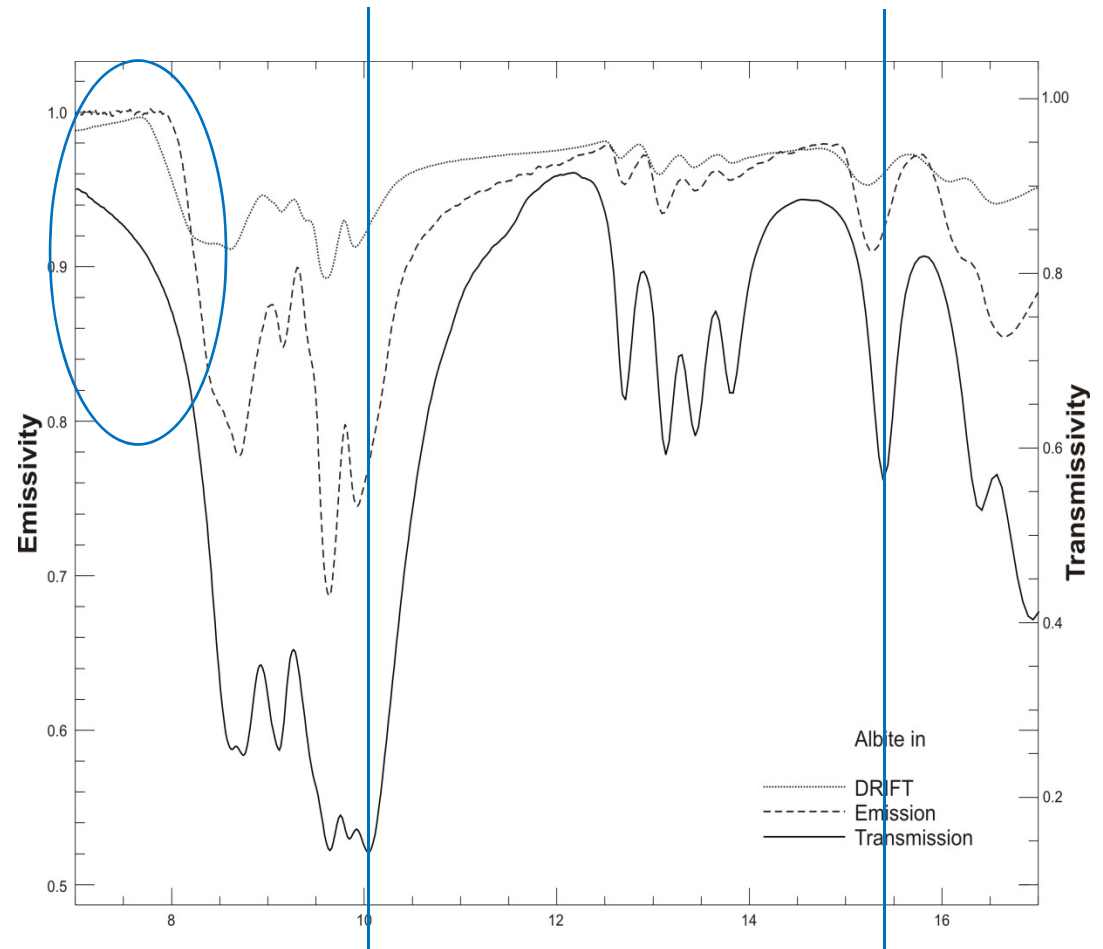
Emission: careful temperature control of sample needed

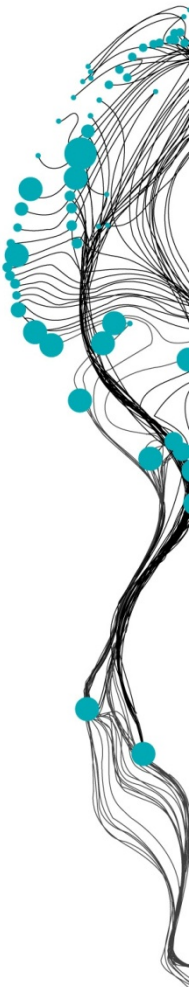


LABORATORY – GEOMETRY CONSIDERATIONS (CONT'D)



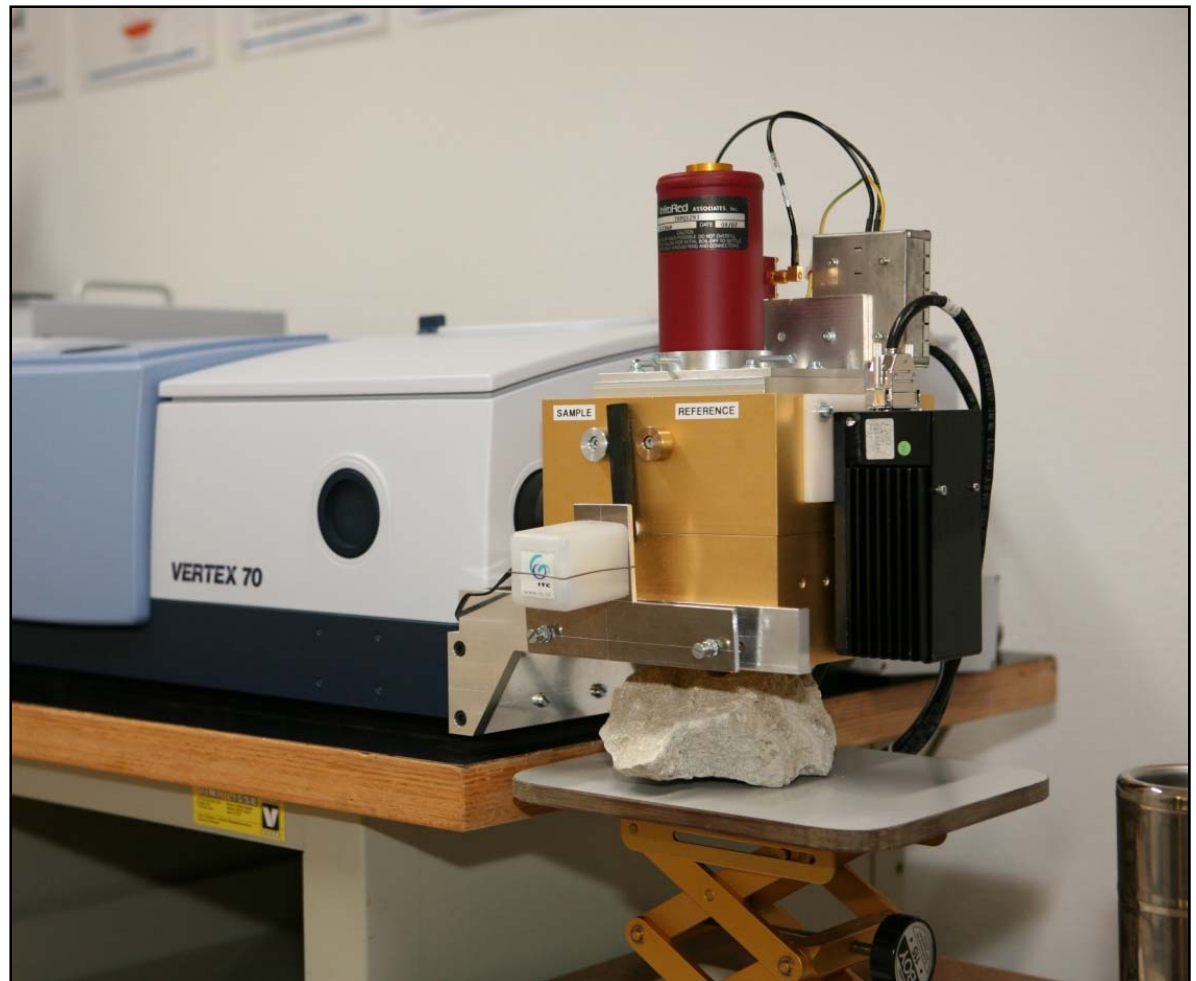
- Same Albite sample
- Measured with DRIFT, DHR, transmission.
- Qualitatively similar
- Quantitatively different (wavelength shifts, relative feature depths ... etc.)

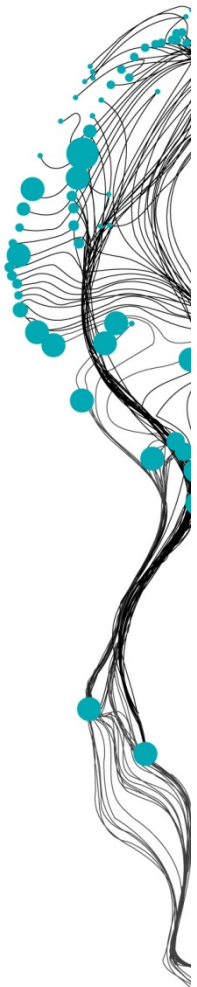




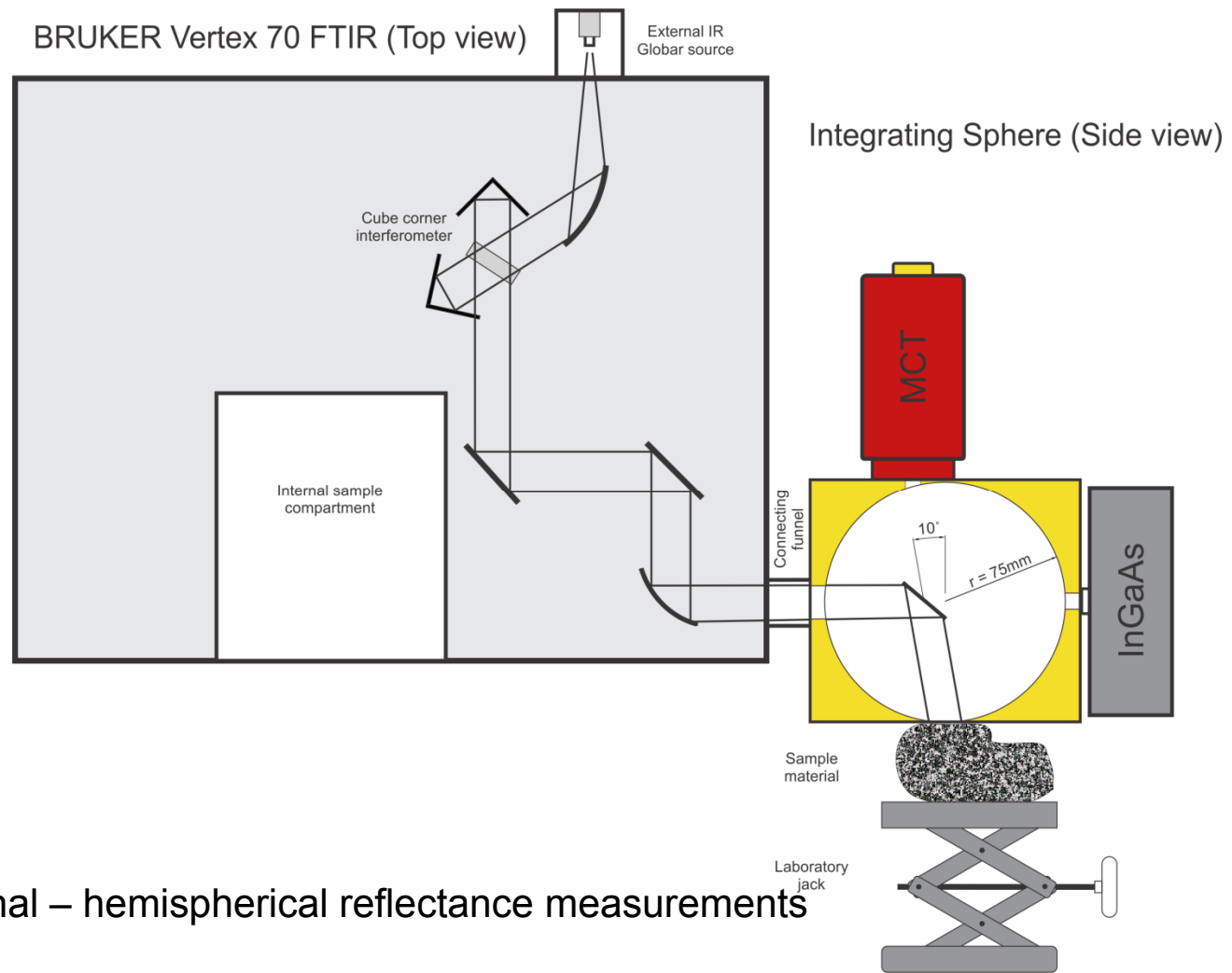
EXTERNAL INTEGRATING SPHERE MEASUREMENT

- Gold integrating sphere
- Double source (SWIR & TIR)
- Double detector (SWIR & TIR)
- Large samples from bottom





EXTERNAL INTEGRATING SPHERE MEASUREMENT (CONT'D)

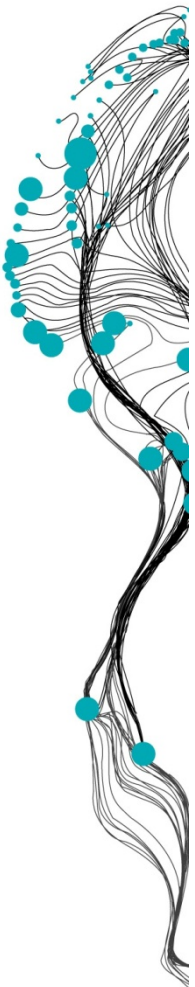


Directional – hemispherical reflectance measurements



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Source: Hecker et al (2011)



EXTERNAL INTEGRATING SPHERE MEASUREMENT (CONT'D)

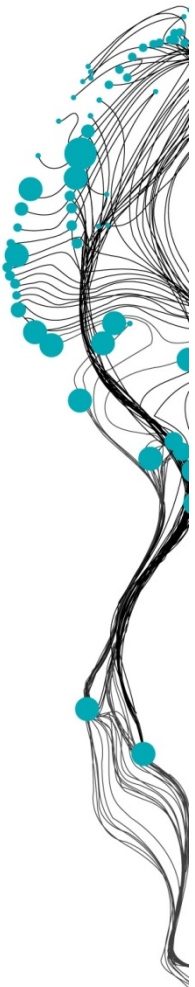


Similar setups at JPL (top left), Geologic Survey Japan (top right) and USGS Reston (bottom center).

Photo credit GSJ: R. Hewson

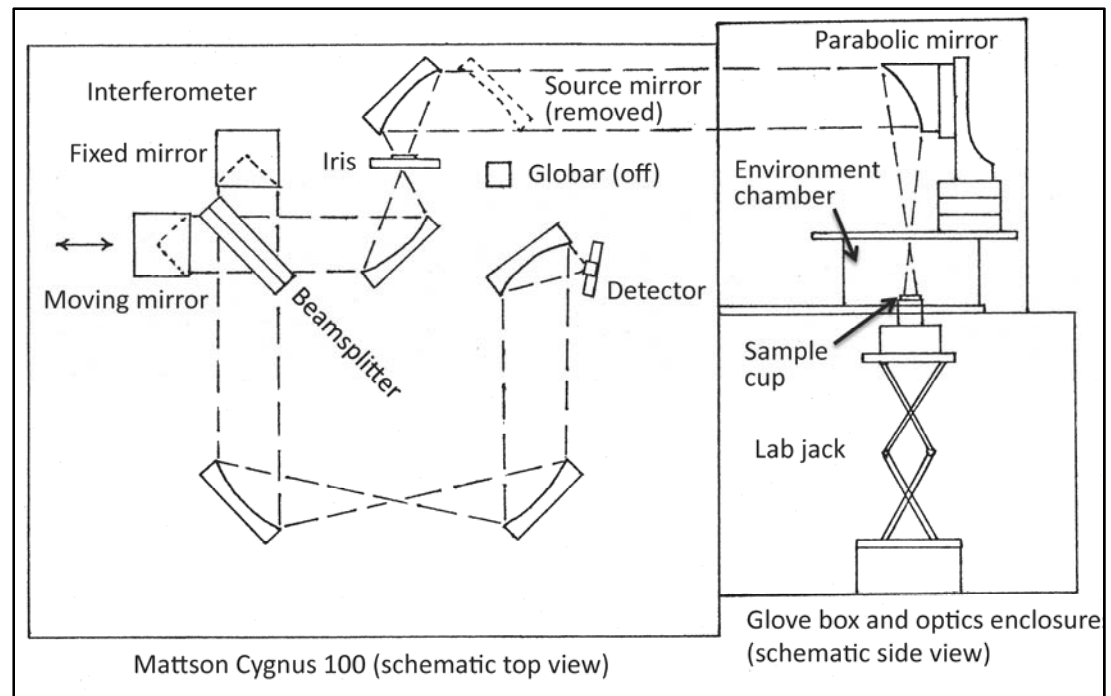


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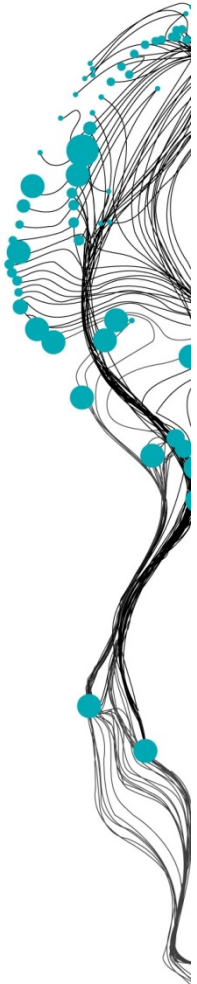


LABORATORY EMISSIVITY MEASUREMENTS

- Schematics of emissive system at Arizona State University
- Planetary community more into emission lab measurements (avoid uncertainty of Kirchhoff law)

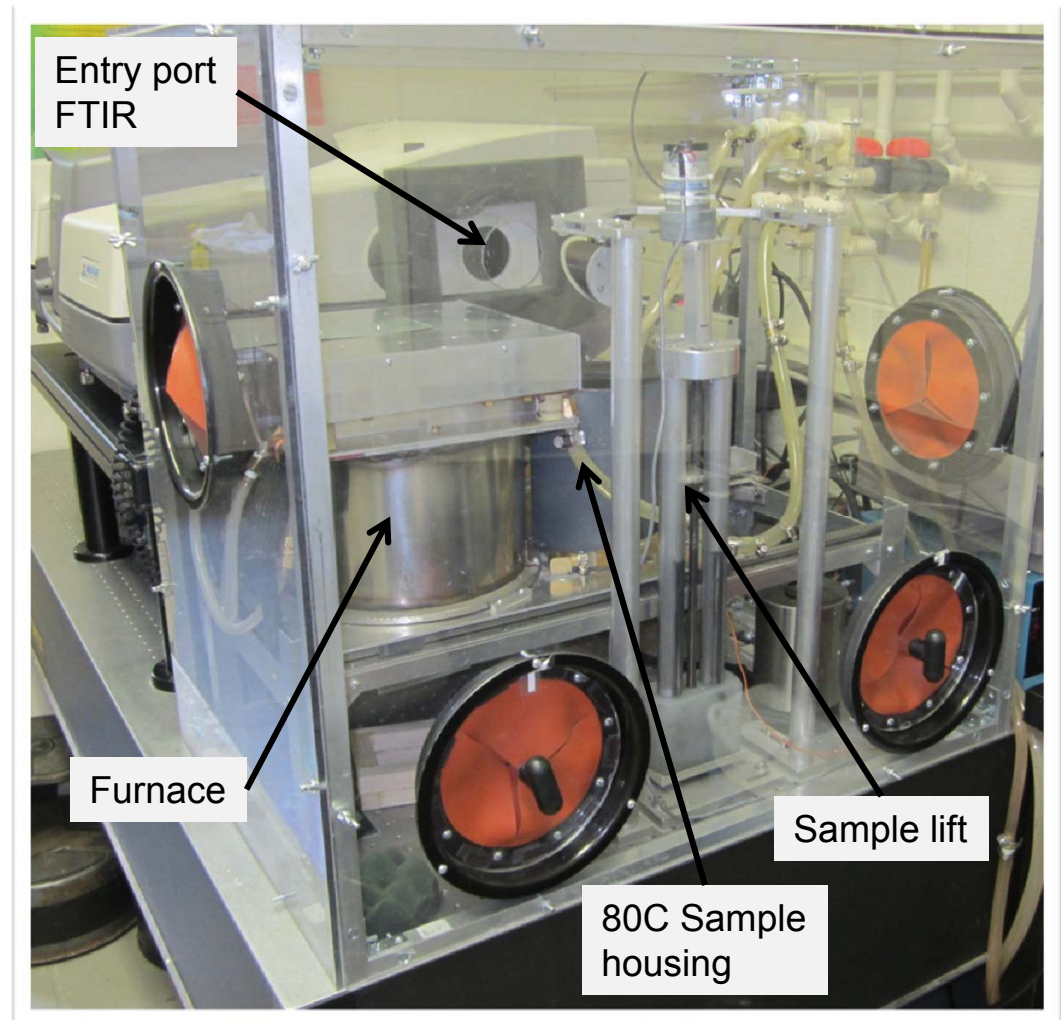


(Schematics: Ruff, 1997)



PITTSBURGH EMISSIVE LAB SPECTROMETER

- Prof. Mike Ramsey (formerly ASU)
- Emission system based on ASU but further developed
- Low temp (80C) and high furnace for high temp (up to 1200C)
- Measurement of emissivity changes when rocks melt





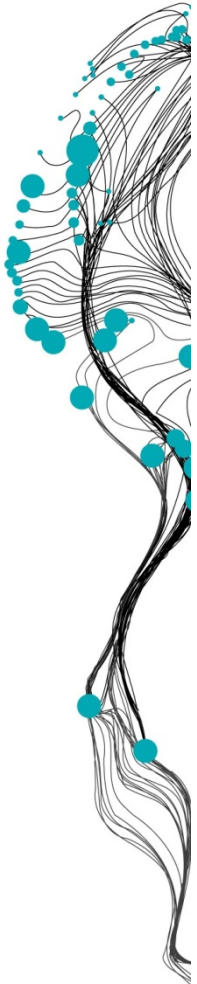
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FIELD – INSTRUMENT CONSIDERATIONS

- Weight
- Power consumption
- Temperature stability
- Ruggedness
- Quality of spectra measured
- ...
- Applicability to Earth Science samples (most systems for open path / gases).



FIELD – STARTING POINT 1 - μ -FTIR

- Pro:
 - Quite light (ca. 7 kg)
 - Low power consumption
 - Designed with ES in mind (down-looking)
 - Ready-to-go system

- Con:
 - Resolution limited
 - Speed of measurement suite
 - Portability OK but not ideal
 - Not rugged
 - Not certified for Europe



Photo source: Richard Bedell, Auex.com



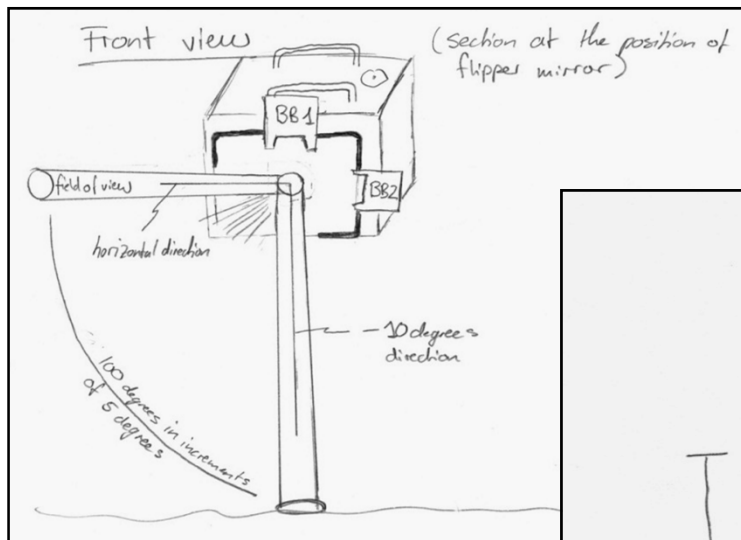
FIELD – STARTING POINT 2 – EMISSION FTIR

- Pro:
 - High resolution
 - Good quality of spectra (high throughput)
 - Rugged

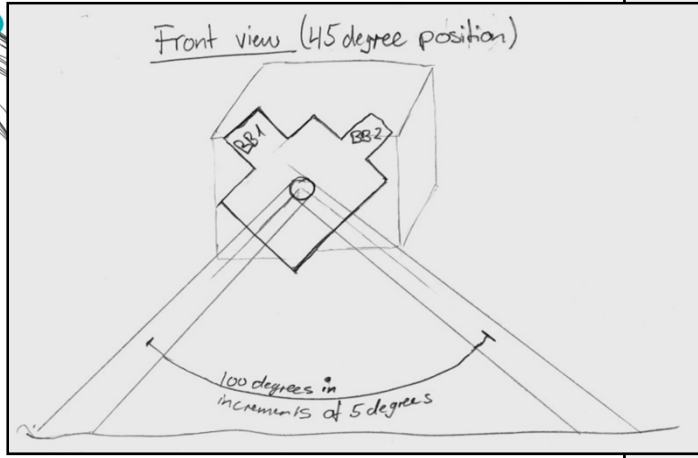
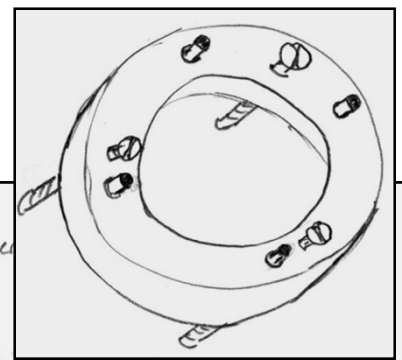
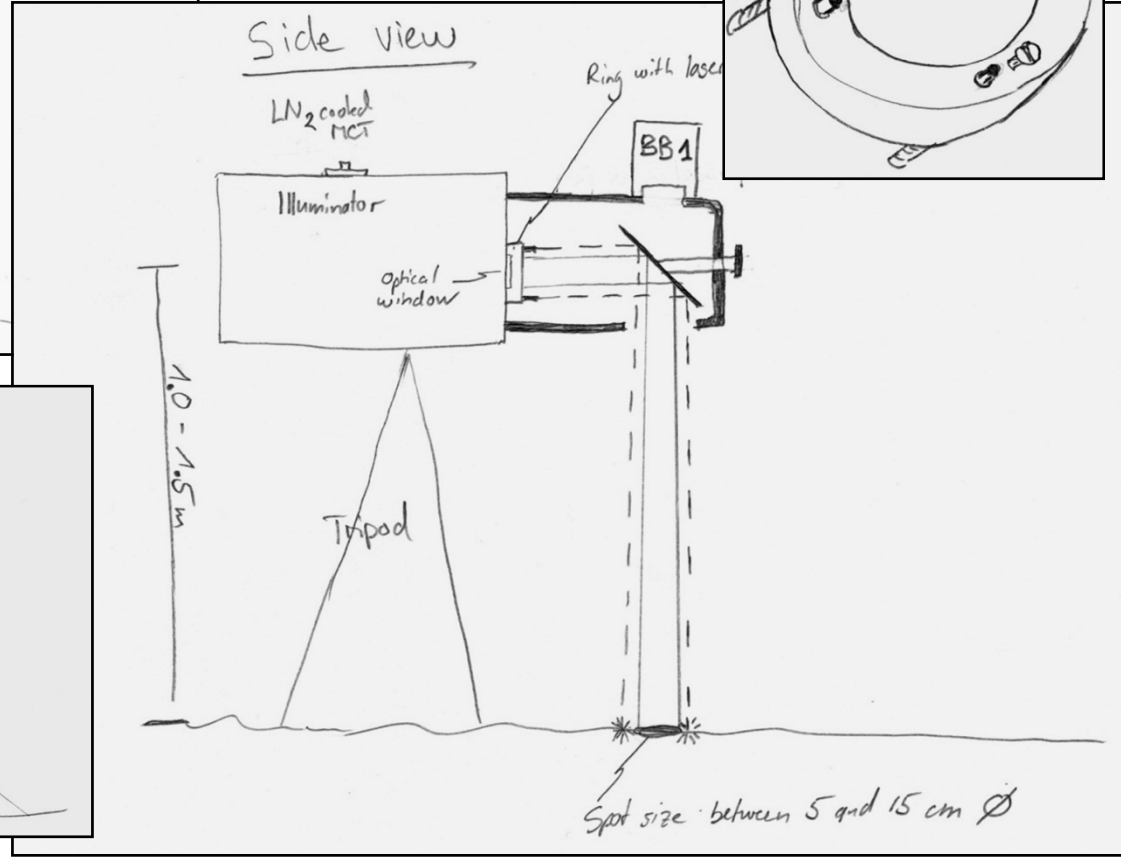
- Con:
 - High power consumption
 - Weight!
 - Made for open path emission measurements. Need specific foreoptics



FIELD - CURRENT ITC SOLUTION

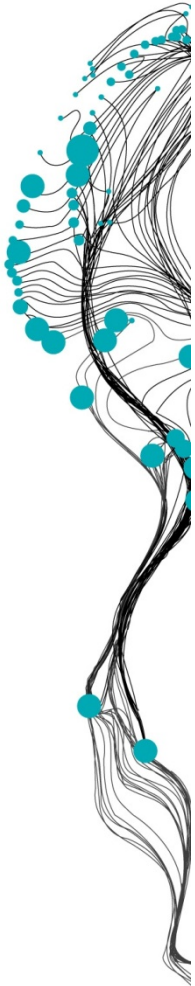


Side view



FIELD – CURRENT ITC SOLUTION (CONT'D)

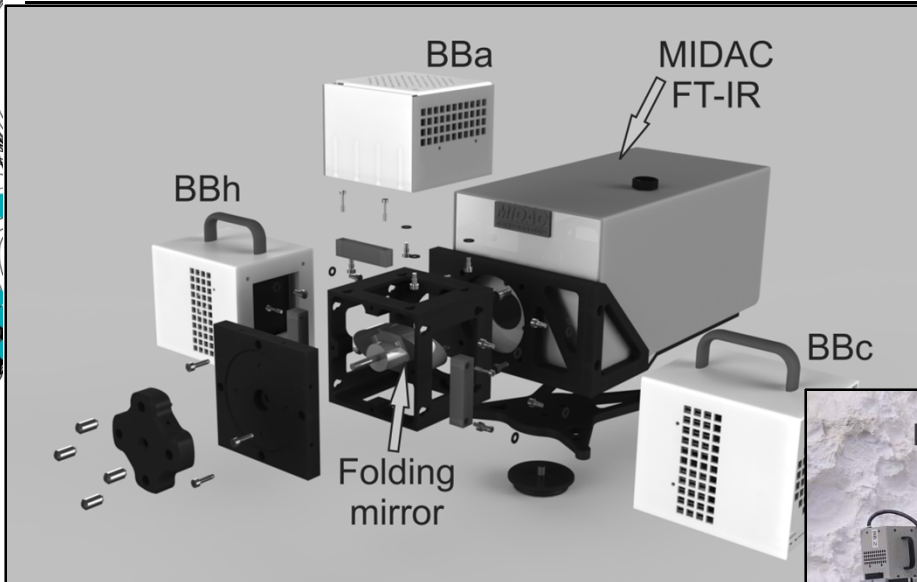
WITH CUSTOMIZED FOREOPTICS



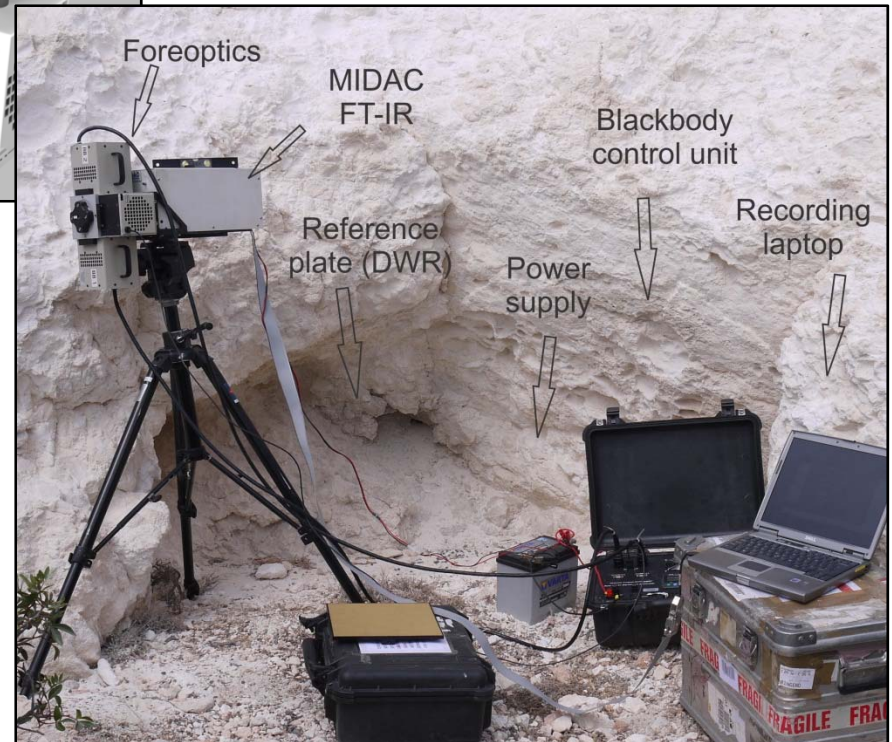
- MIDAC Illuminator M4401
- Non-hygroscopic ZnSe optics
- Heavy duty, sealed cast aluminium housing (ca 15kg)
- IN2 cooled



FOREOPTICS – MARK2



- Next development stage of foreoptics by NERC-FSF.



Photot credit: Chris MacLellan, University of Edinburgh



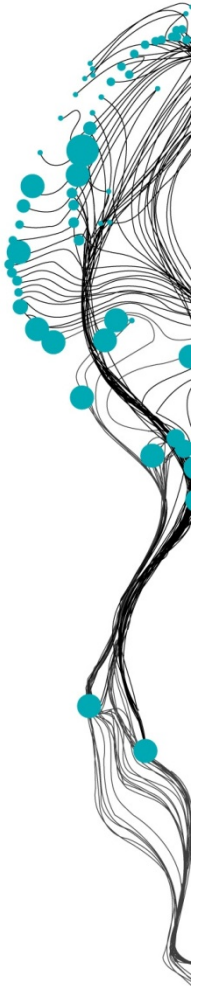
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FIELD IMAGING SPECTROMETER – TELOPS HYPERCAM

- Emissive system
- Imaging FTIR
- Spectral range: 7.7 - 11.5 μm
- Image pixels: 320 x 256
- Calibration: 2 Blackbodies
- Weight: ~30 Kg





AGILENT EXOSCAN 4100

- Diffuse reflectance measurements
- Not quantitatively comparable to emissive systems
- Lightweight: ~3 Kg
- Comparable in use to PXRF





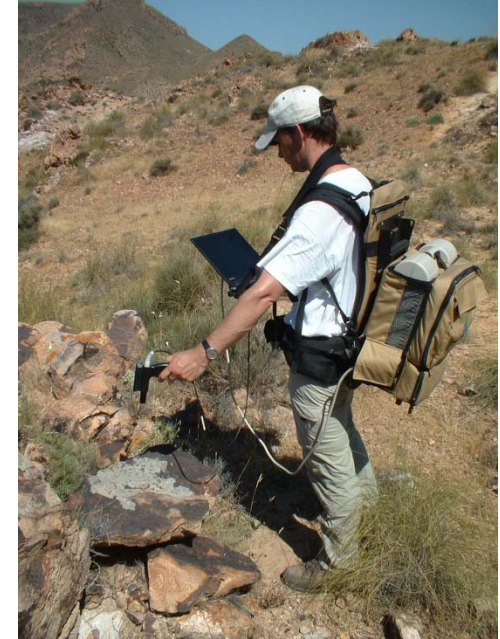
CONSIDERATIONS - LABORATORY

- **Decide: speed or absolute emissivity values?**
 - Speed: bi-directional (cheap, fast, high SNR)
 - Abs. Emiss: more effort, costs, measurement time
- **Abs. Emissivity:**
 - Sphere: long measurements, costs (1kEUR per cm diameter)
 - Emission: sample temperature control!



CONSIDERATIONS - FIELD

- TIR field instruments not in ASD-like category
- Decide: 50 kg equipment to field or 50 kg samples to lab?
- Personal take on this question: Bring samples to lab except:
 - Calibration Airborne campaigns
 - Vegetation (?Lichen)
 - Extremely large samples (e.g. entire quarry wall)
 - Undisturbed soils and evaporite crusts (sometimes sample rings possible?)





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MAPPING METHODS

- TIR preprocessing fundamentally different (e.g. T ϵ S)
- After reduction to ground leaving radiance, same hyperspectral tools as VNIR-SWIR mineral mapping, e.g.:
 - Linear unmixing
 - Partial Least Squares Regression (PLSR)
 - Mixture Tuned Match Filter (MTMF)
 - Spectral Angle Mapper (SAM)
 - Feature fitting
 - ...
 - etc.

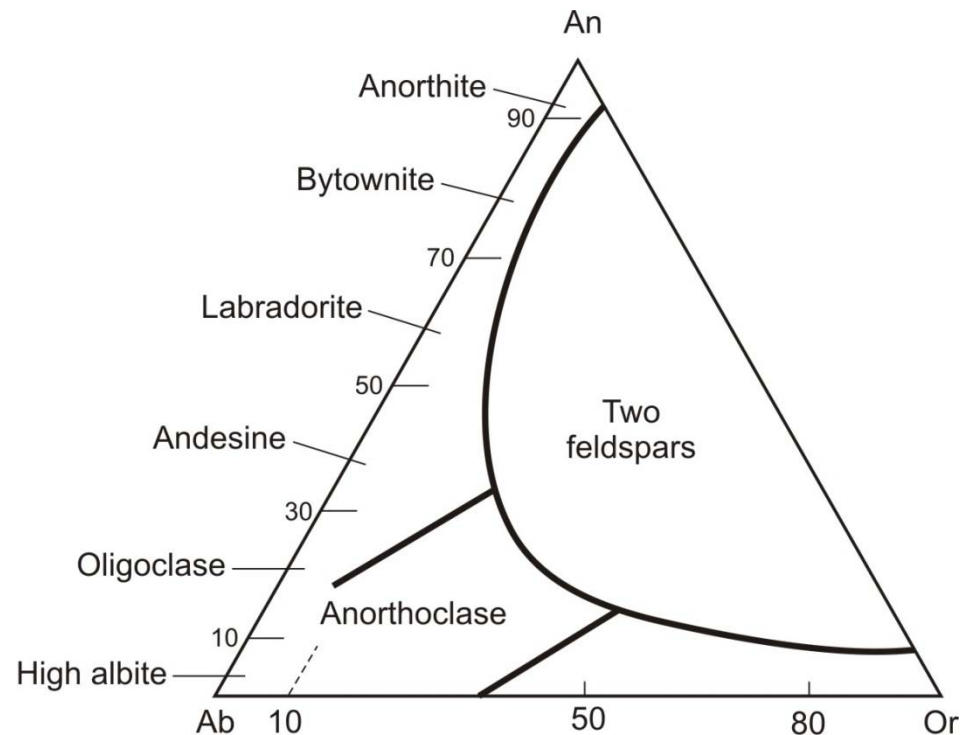


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CASE STUDY: TIR – PLSR – MINERAL MODES

- Determining quantitative rock compositions (= “mineral modes”) on lab and airborne TIR spectra.
- Quartz modes
- Alkali feldspar and plagioclase modes
- Plagioclase compositions



Source: Hecker et al (2010)

QUANTITATIVE TIR SPECTROSCOPY

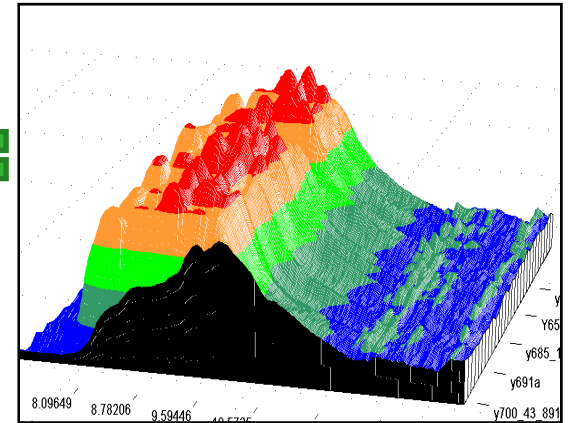
LINKING SPECTRA TO MINERALOGY AND MINERAL CHEMISTRY



Thin section blocks



Bruker FTIR



TIR spectra

Microscopic description:

Percent	Mineral	Size/shape	Diagnostic Features
~(55)53	plagioclase	~.5-1 / anhedral	Et+ L's = 0%, 0-4%
7	quartz	~.5 mm / anhedral	
20	epidote	~.05-1 / anhedral	B ₂ V ~ 75-80°; 39.2° bf; non-pleuroclastic
4	chlorite	~.05-1 / anhedral	brown bf (anom); length fast; ga → pgn. (locally v. white st. sil & w. or claus)
10%	Ksp (parthite)	~.5-2 mm / anhedral	B ₂ V ~ 60°
3	calcite	~.05-1 / anhedral	
2	sphene	~.1-1.5 / anhedral	brown
tr	zircon	~.05 mm / anhedral	(3,5)
44-42	apatite	~.05 mm / anhedral	
1/2	goethite		
(18)	Mfcs		

opaque (Rutile? or Mt?) to reflect Mt!

Textural and structural features: paragenetic relations:

Ksp = 10%^{20%} perthite (granitic; cleavag. controlled!); locally 10% sil & to epidote.
Ksp v. interstitial, poikilitic; milled appearance suggests Ksp → Al?!

Plag → Epidote (5-30%) ± calcite 0.5%
unzoned, An = 45-22

sampleloc	lithunit_shpft	TS_name_long	samplelv	lithunit_TSsheet	alteration	quartz	plagioclase	Ksp	diopside	clinoclase	garnet	plagioclase	Bl_1gn	Bl_2nd	Bl_total
5630nd	6513_90067_26	Y-563	OMD	OMD	pervasive S-2 or ES-2	17	70	0	8	35	0	0	0	0	0
6020jamp	6513	Y-6513	OMP	OMP	wk ab-chl	3	19	4	0	0	0	35	0	2	2.5
6513nd2	6511_16_891	Y-6511	Jamp	?		5	30	16	0	0	0	0	0	0	10
4640jamp	446_90067_5	Y-448	OP-dike	di	chl rt ep (v)	24	25	24	0	0	24	0	0	0	0
7878jamp	787_11_728	Y-787	70qpmj	?	wk clay	30	35	27	0	0	0	25	0	0	5
684jamp	684_40_237	Y-684	Jamp	?	wk serotite	32	27	36	0	0	0	8	0.5	0	0.5
684jamp	684_41_237	Y-684	Jamp	?	wk potassic/ep-chl	5	21	7	0	0	0	0	0	1	1
684jamp	684_18_891	Y-688	Op monozonitite	?	Op monozonitite	0.5	8	0	0	0	0	0	0	1	1.25
684jamp	688_18_891	Y-688	OMP	?	Op	20	25	28	0	0	0	22	0	0	5
690b	690b_42_237	Y-690b	Jamp	?	wk chl py ser / superg clay	3	30	3	0	0	0	33	0	0	0
691a	691a	Y-691a	Jamp	?	transitional phase of border of monzonite	22	32	31	0	0	0	35	0.5tr	0.25	0.25
700b-jamp	700_43_891	Y-700	Jamp	?	wk Ksp-clay-ep	35	20	33	0	0	0	29	1	0	1
708b	708b_50_237	Y-708b	Jamp	?	wk potassic	10	50	15	0	0	0	0	0	10	10
700jamp	770_100-3_1nk	Y-770	PL Lake Hill	Brsh		25	35	25	0	0	0	17	6	1	7
660nd	660	Y-660	OMP	wk SW		10	40	8	0	0	0	20	3	0	3
680nd	689	Y-689	OMP	Na-Ca		12	45	9	0	0	0	28	3	0	3
690a	690	Y-690a	OMP	wk Na-Ca		7	42	12	0	0	0	0	0	0	0
691a	691a	Y-691a	OG	PA		30	22	32	0	0	0	24	6	0	6
691b	691b	Y-691b	OMP	wk PA		6	55	6	0	0	0	7	0	0	0
464jamp	446_90067_4	Y-446	OP-dike	Phlog-Chl-ep		0.25	25	0	0	0	0	36	0	0	0
320nd	320_90067_11	Y-320	OMP	Ac / ep		9	50	9	0	0	0	0	0	0	0
321nd	321_90067_12	Y-321	OMP	wk ES		10	55	12	5	0	0	25	3	0	3
322nd	322_90067_13	Y-322a	OMP	wk PA		29	26	29	0	0	0	24	6	0	8
323nd	323b_90067_14	Y-323b	Andeste Dike	?		1	65	0	0	0	0	0	0	0	0
335nd	335_grinding_b	Y-335	Andeste; Arthesia?			0	65	0	0	0	0	10	0	0	0
680b	680b_13_237	Y-680b	Jamp	?		1	24	0	0	0	0	0	29	0	0.25
691d	691d_41_237	Y-691d	Jamp(j)	Op-Tm-Ser		32	21	27	0	0	0	31	0.125	4.5	4.625
797nd	797_15_7	Y-797	Artesia Fm	wk Ep clay-Ab		0	53	0	0	0	0	0	0	0	0

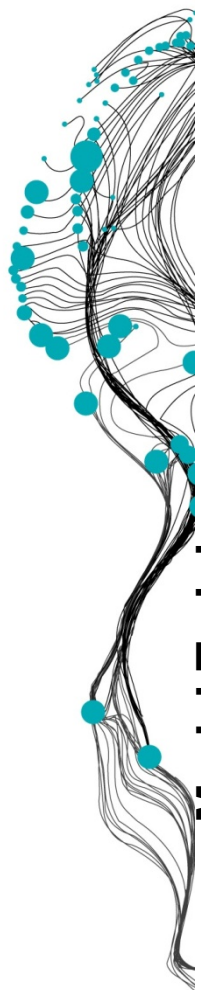
TS descr. In Spreadsheet



Thin section description
UNIVERSITY OF TWENTE.

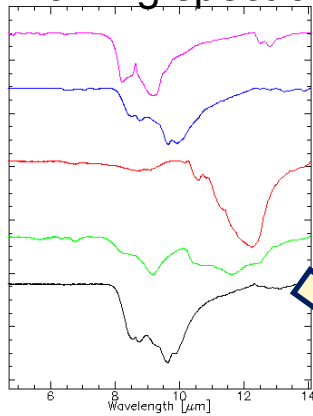
QUANTITATIVE TIR SPECTROSCOPY

LINKING SPECTRA TO MINERALOGY AND MINERAL CHEMISTRY



Model Training

Training spectra

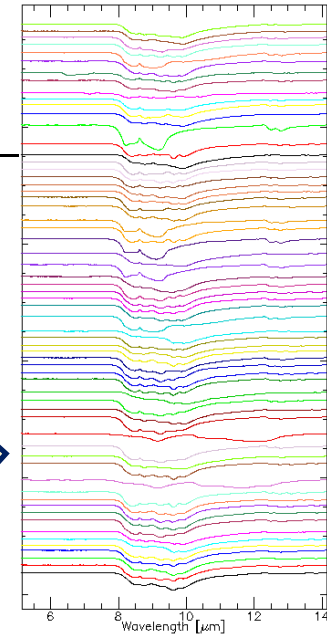


Reference methods

XRD
XRF
SEMP
Thinsection

PLSR
Model

Sample spectra



Predicted
mineralogy
and
mineral
chemistry

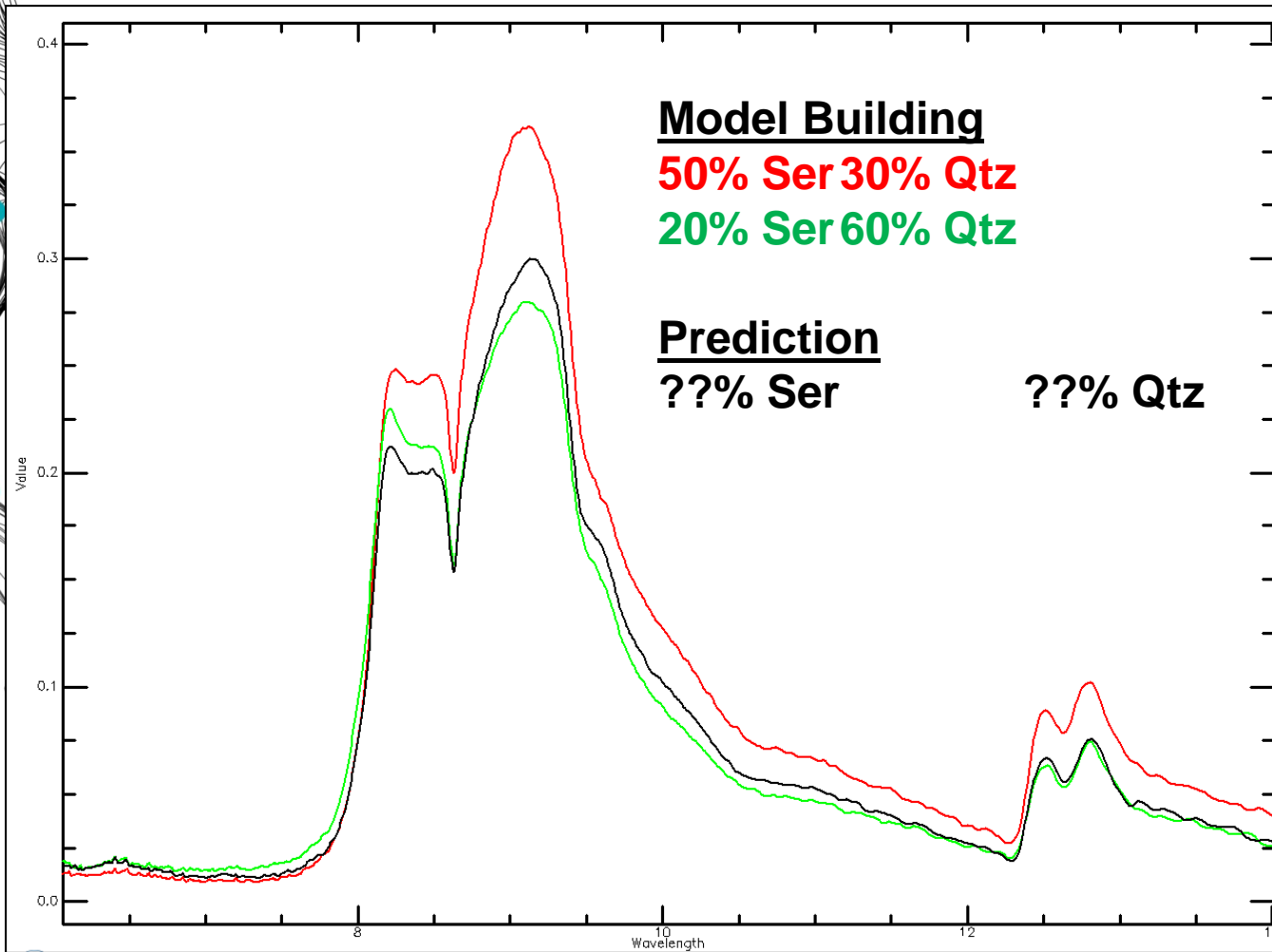
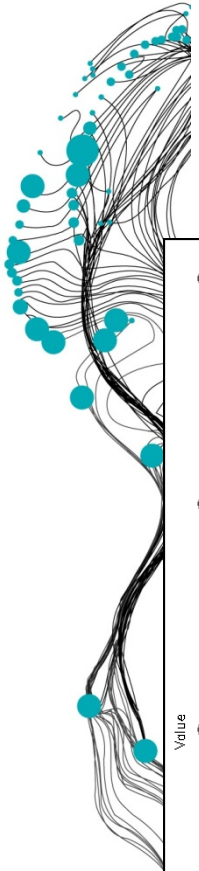
Model Prediction



WHAT IS PLS?

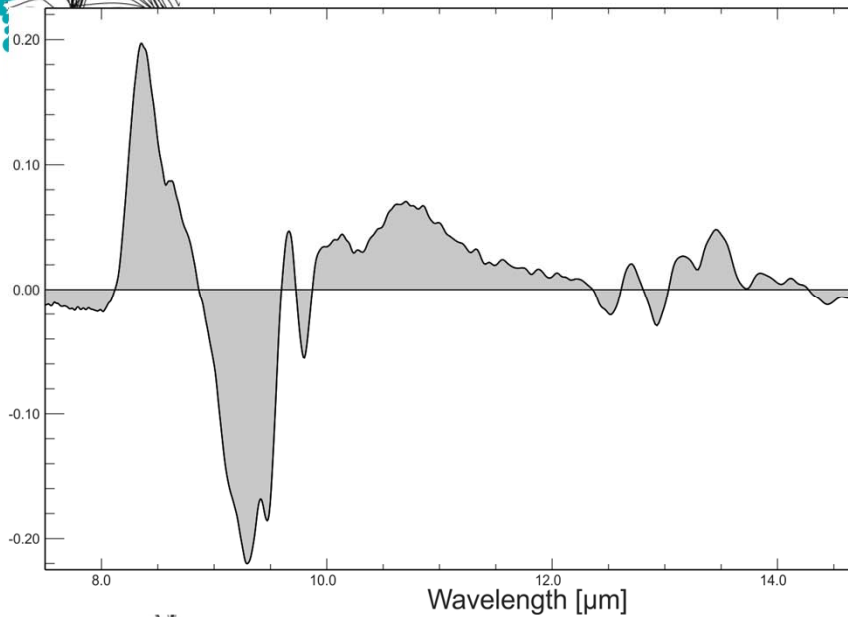
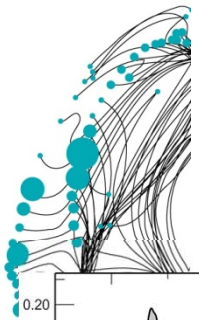
- Regression method
- Links attribute data to spectra
- Decomposes into components similar to Principal Component Analysis
- Good for spectroscopy:
 - Compresses info into a few components
 - Can deal with lots of bands and selects the most important
 - Deals well with correlated attributes (adjacent bands often 99% correlated)

SIMPLIFIED PLS EXAMPLE

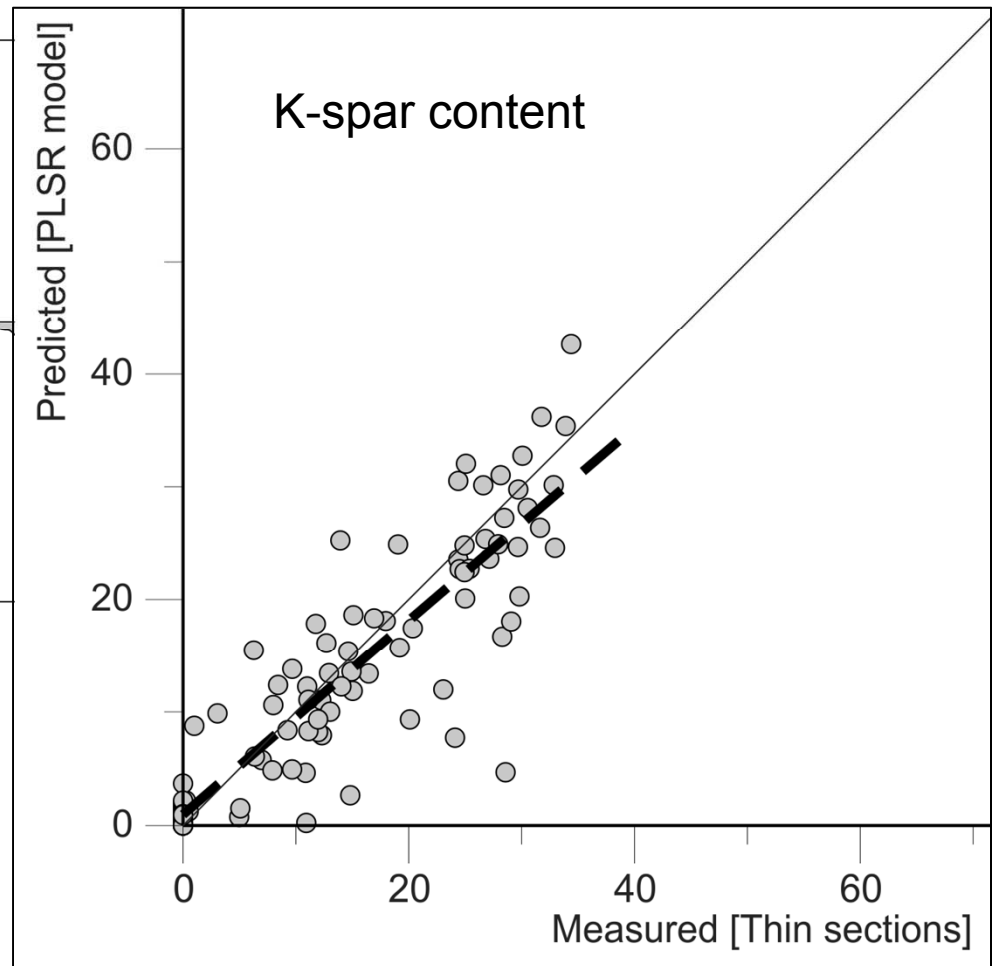


PLSR ON TIR SPECTROSCOPY

PREDICTION RESULTS

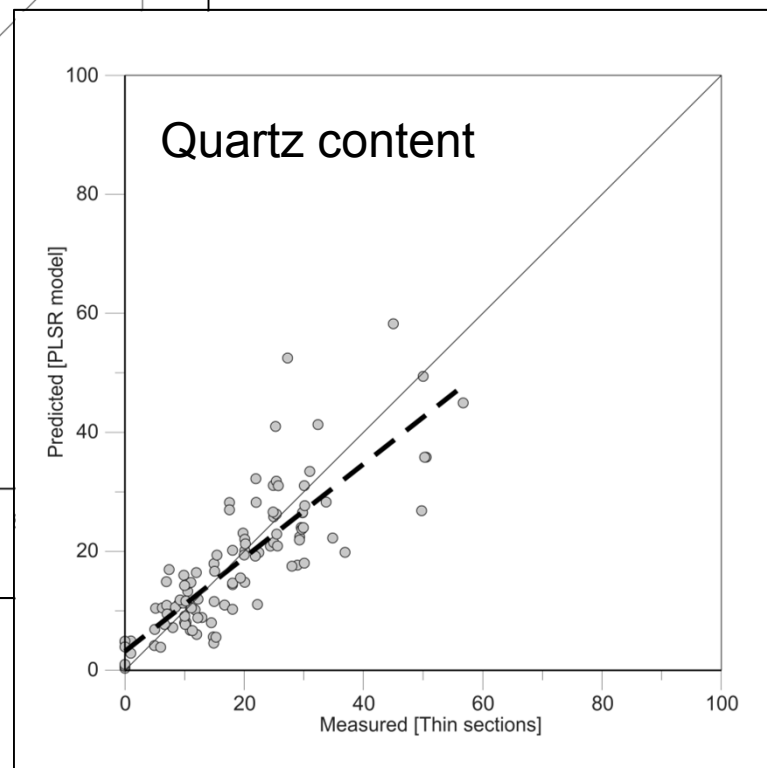
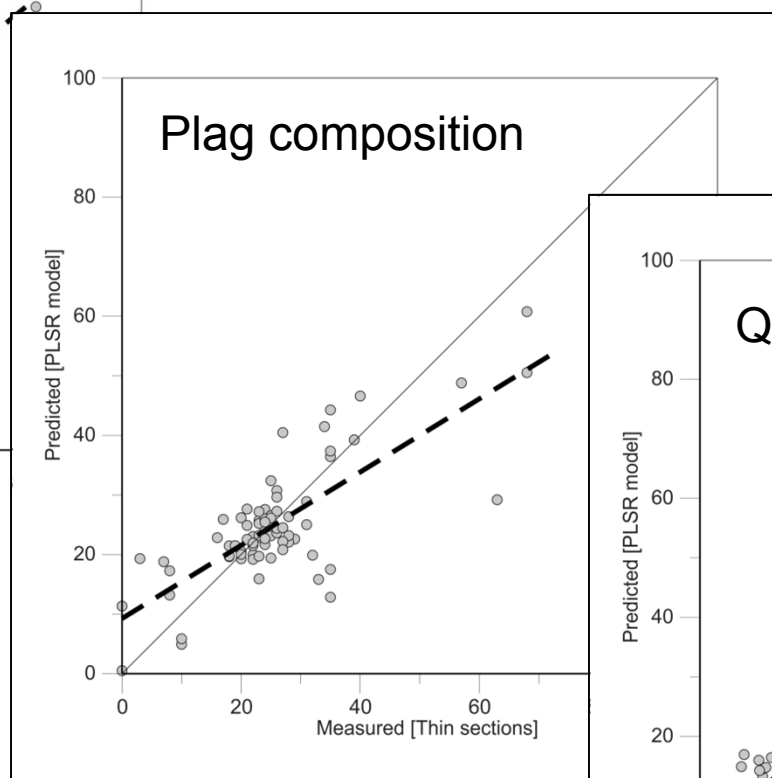
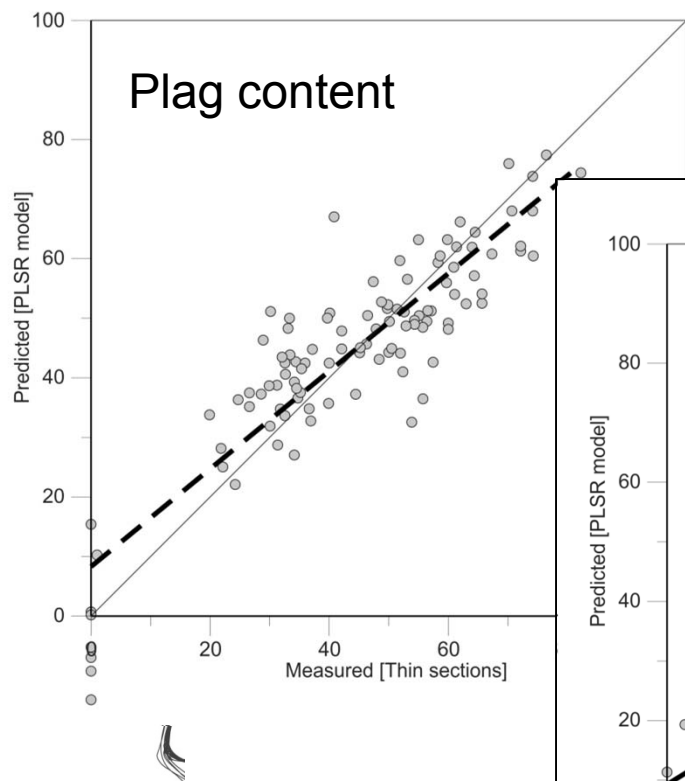


- Regression coefficients and meas. vs. predicted plot





PLSR ON TIR SPECTROSCOPY PREDICTION RESULTS



Source: Hecker et al (2011)

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PLSR ON TIR SPECTROSCOPY

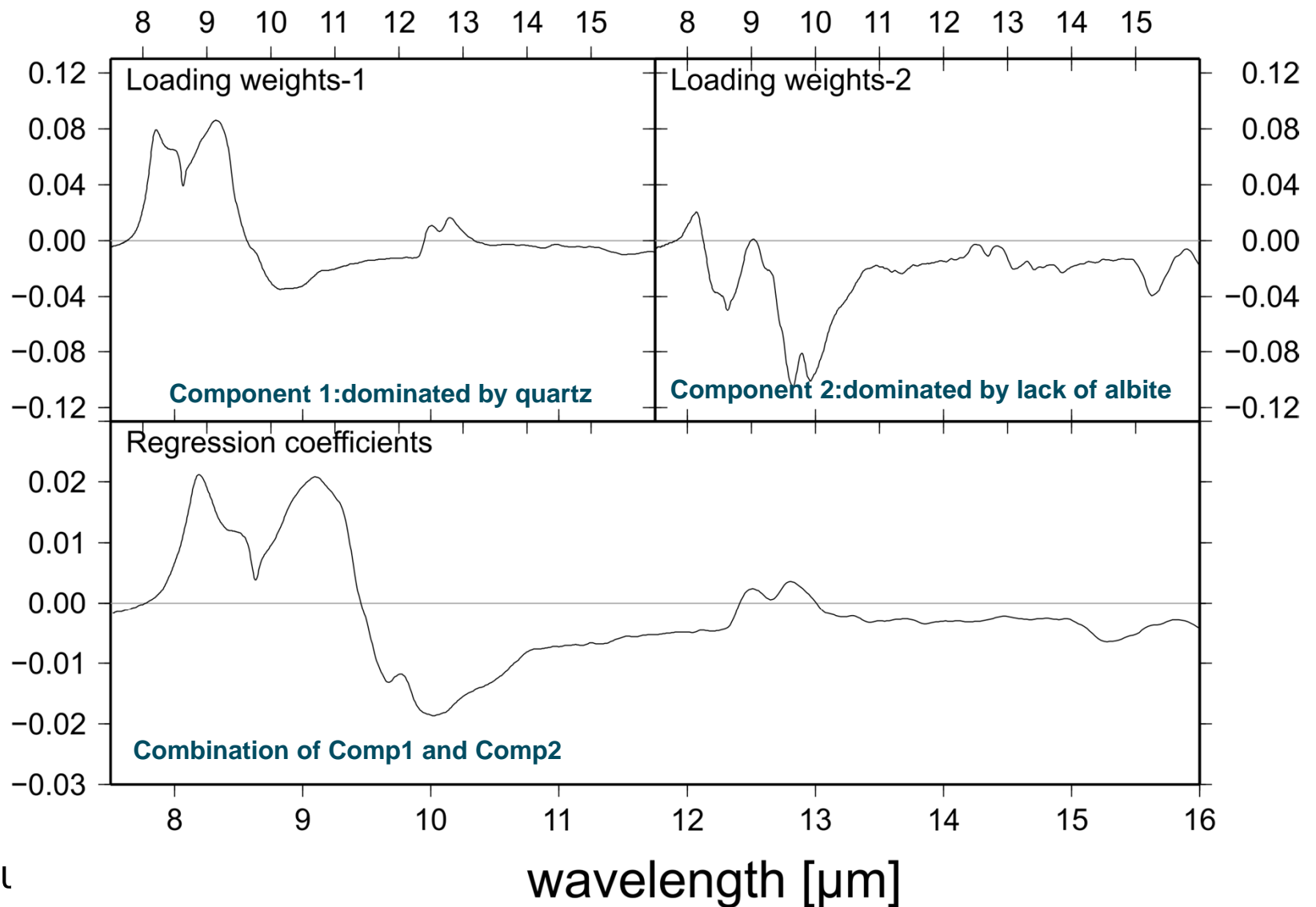
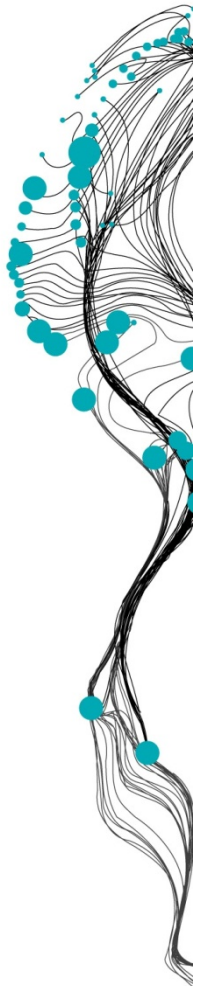
PREDICTION RESULTS



Mineral	Ksp	Plg	Qtz	Plgcomp
Number of LV's used	5	4	2	5
RMSEP [in %abs]	5.13	8.52	6.90	7.79
R2	0.81	0.80	0.70	0.59
slope (of regression line)	0.86	0.82	0.79	0.61

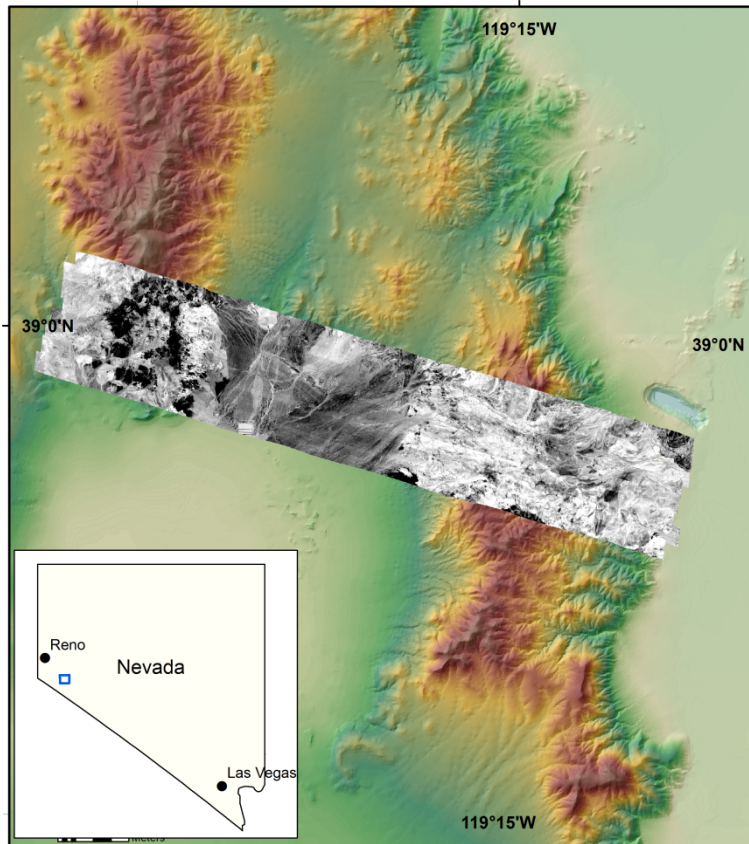
PLSR ON TIR SPECTROSCOPY

MODEL INTERPRETATION (CASE OF QUARTZ)



APPLICATION TO AIRBORNE HYPERSPECTRAL IMAGE DATA

AEROSPACE CORPORATION'S SEBASS SENSOR OVER YERINGTON BATHOLITH, NEVADA



Data courtesy Dr. Dean Riley, Aero.org



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Source: <http://www.lpi.usra.edu/science/kirkland/Mesa/text.html>

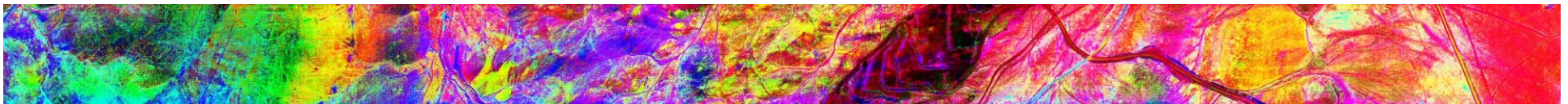
YERINGTON FIELD IMPRESSIONS



Yerington Mine (porphyry Cu)



MacArthur Mine (porphyry Cu)



SEBASS d-stretched Colour Composite

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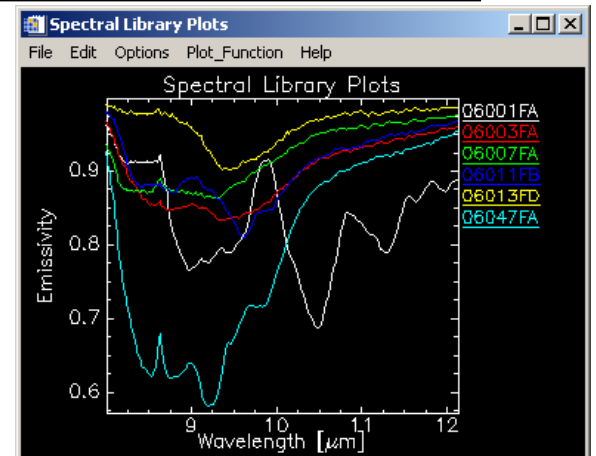
YERINGTON FIELD IMPRESSIONS (CONT'D)

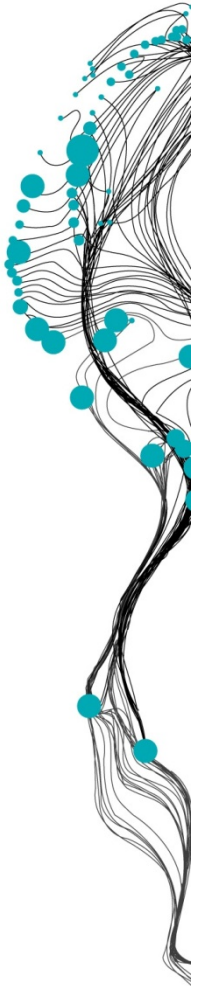


Breccia with
Cu-Oxides



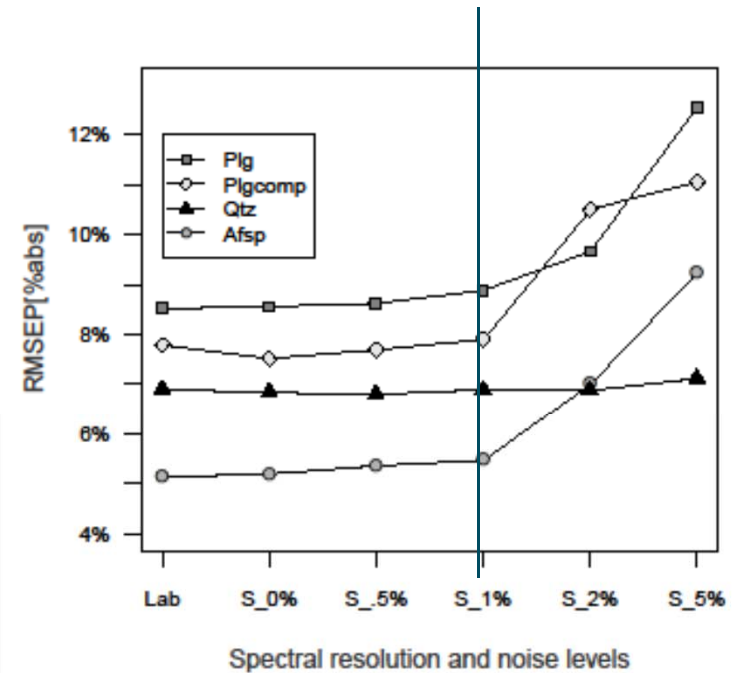
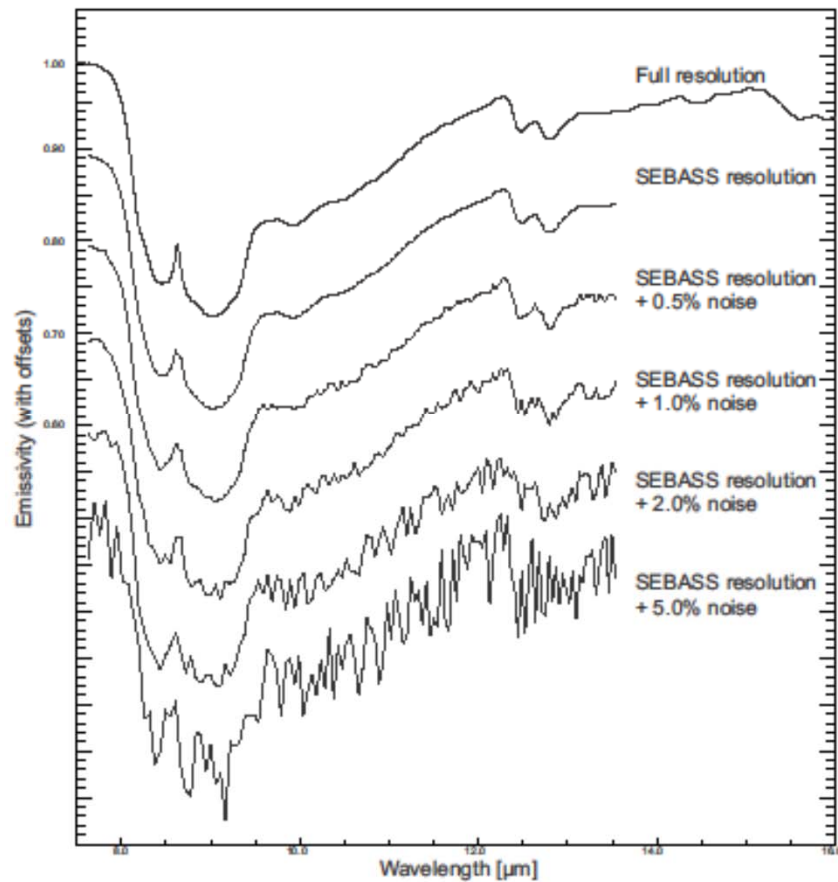
Granite w/ Epidote
and Hornblende

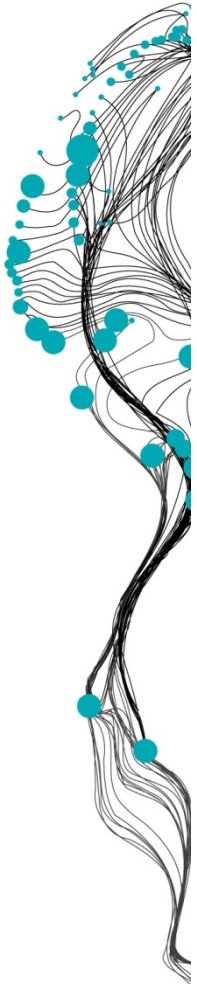




MODELING FIRST

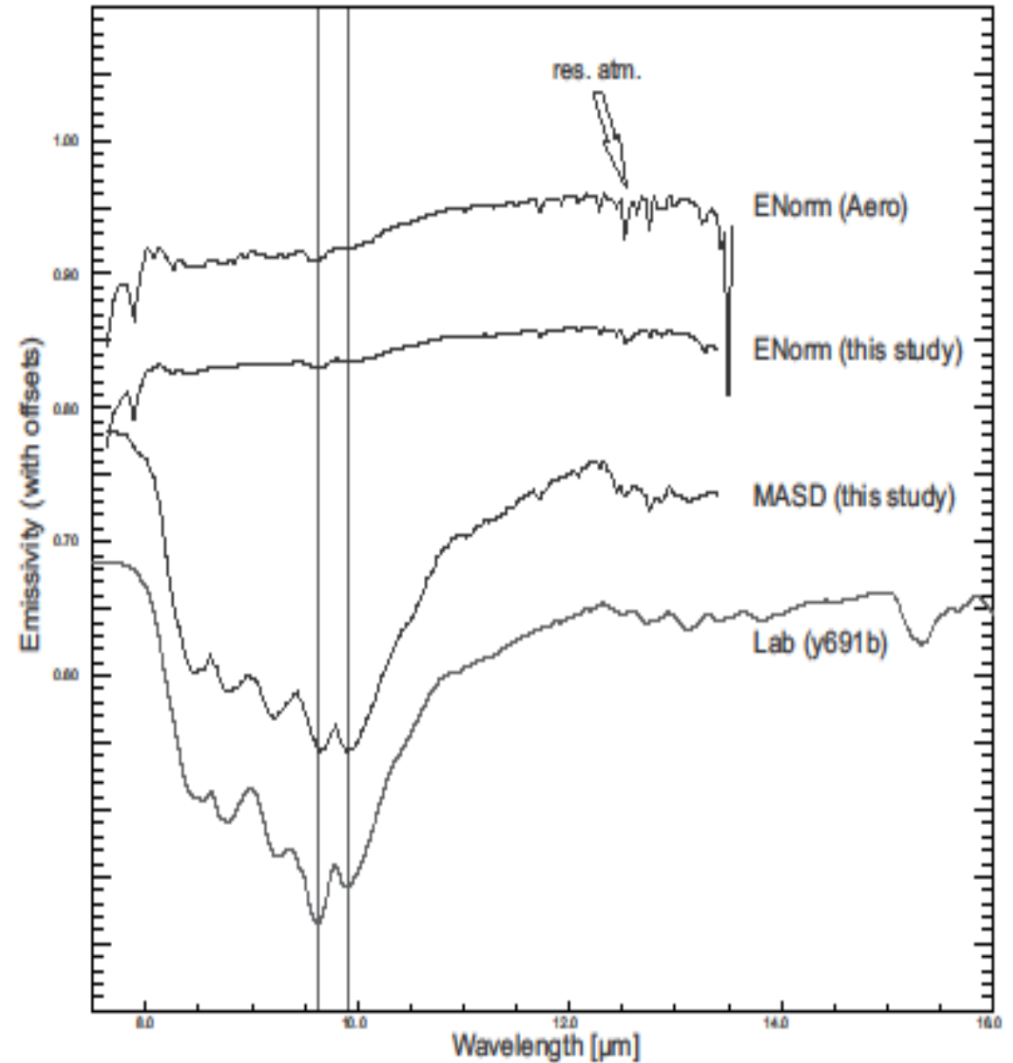
- Adding noise up to 1% (absolute) to emissivity spectra gives OK results





NORMALIZING SPECTRAL CONTRAST

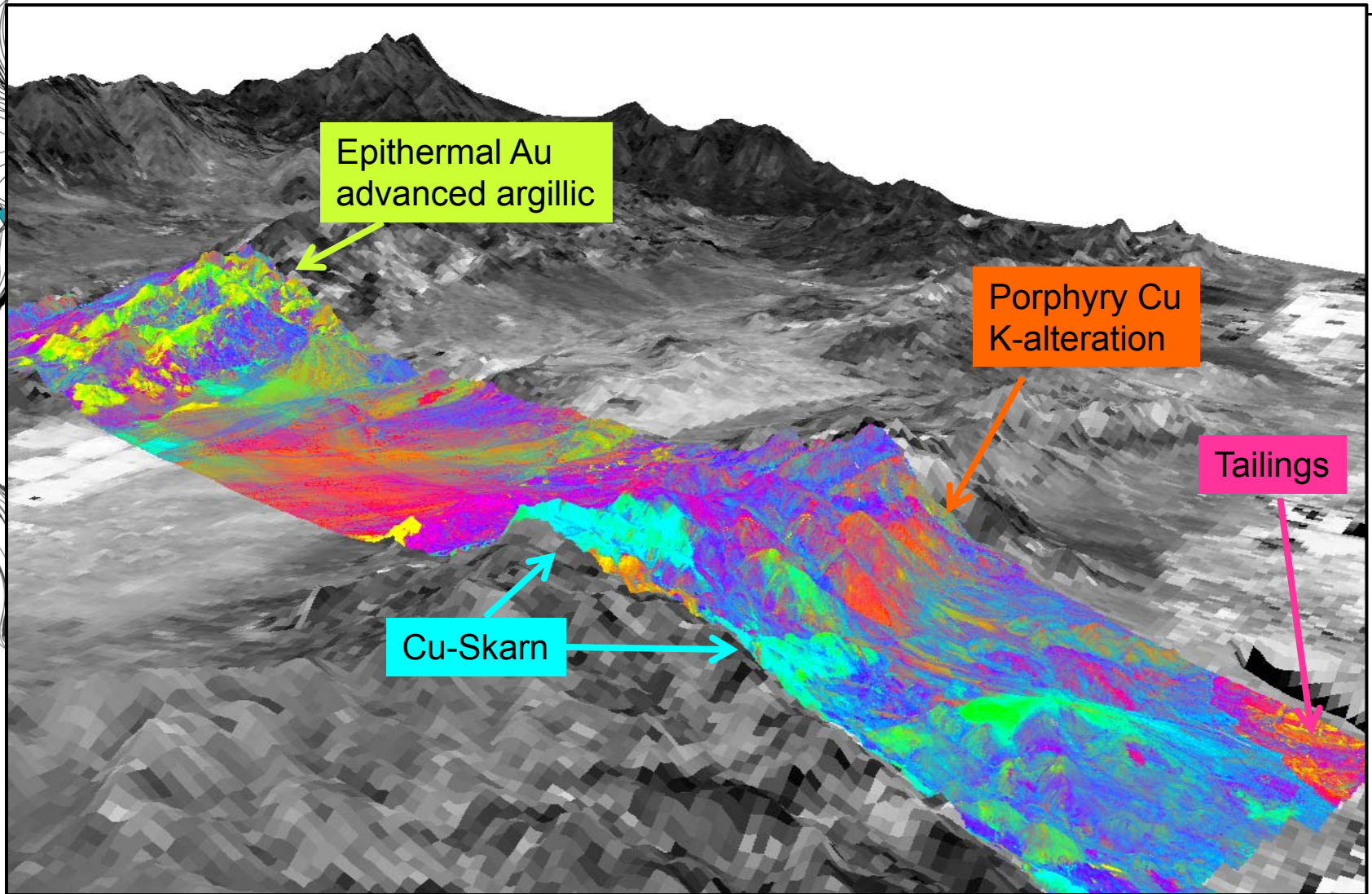
- Airborne spectra have minimal spectral contrast
- For quantitative results, spectral contrast needed normalization.



Source: Hecker (2012)

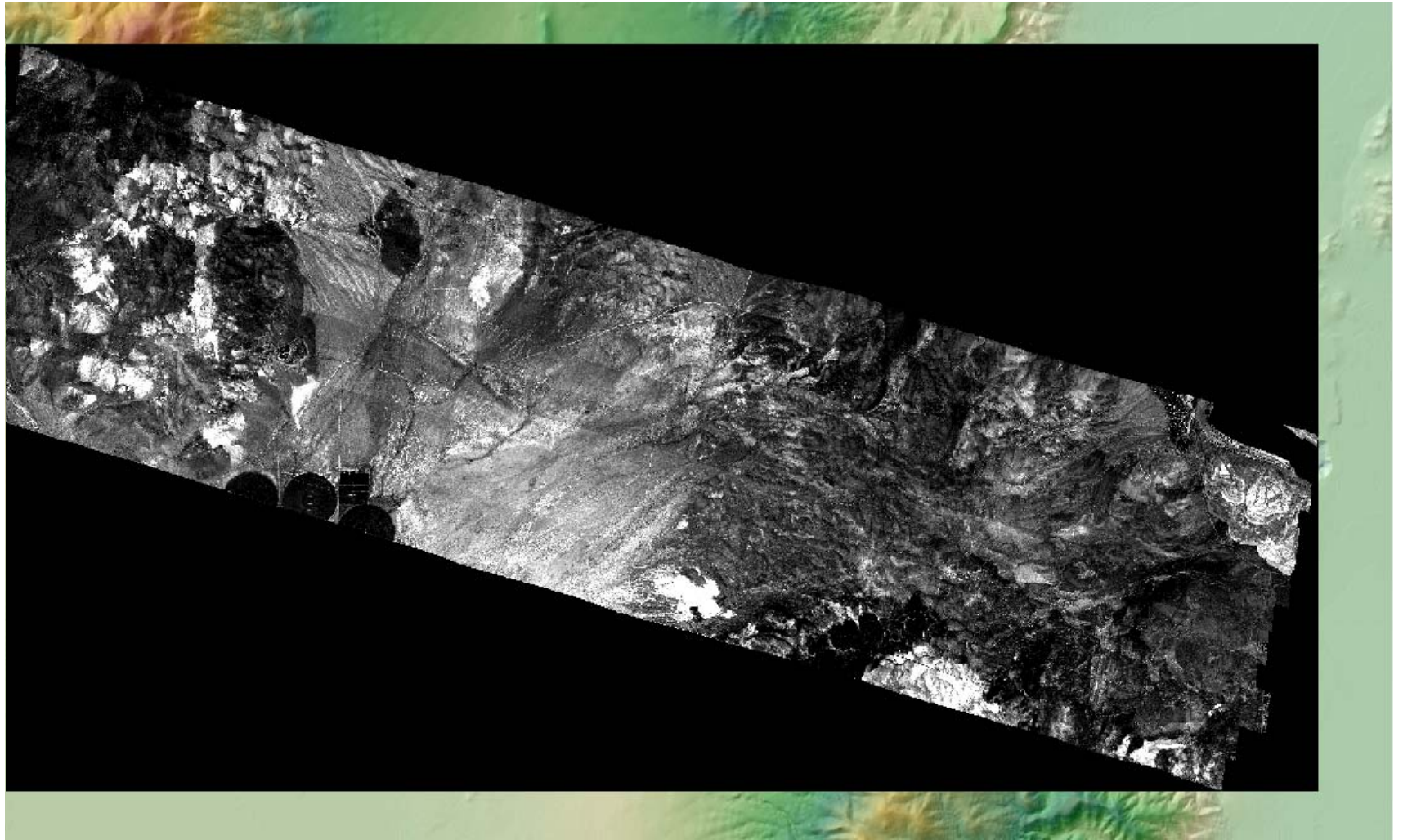
AIRBORNE HYPERSPECTRAL IMAGING

YERINGTON TIR COLOUR COMPOSITE RGB = (11.1, 9.64, 9.06)



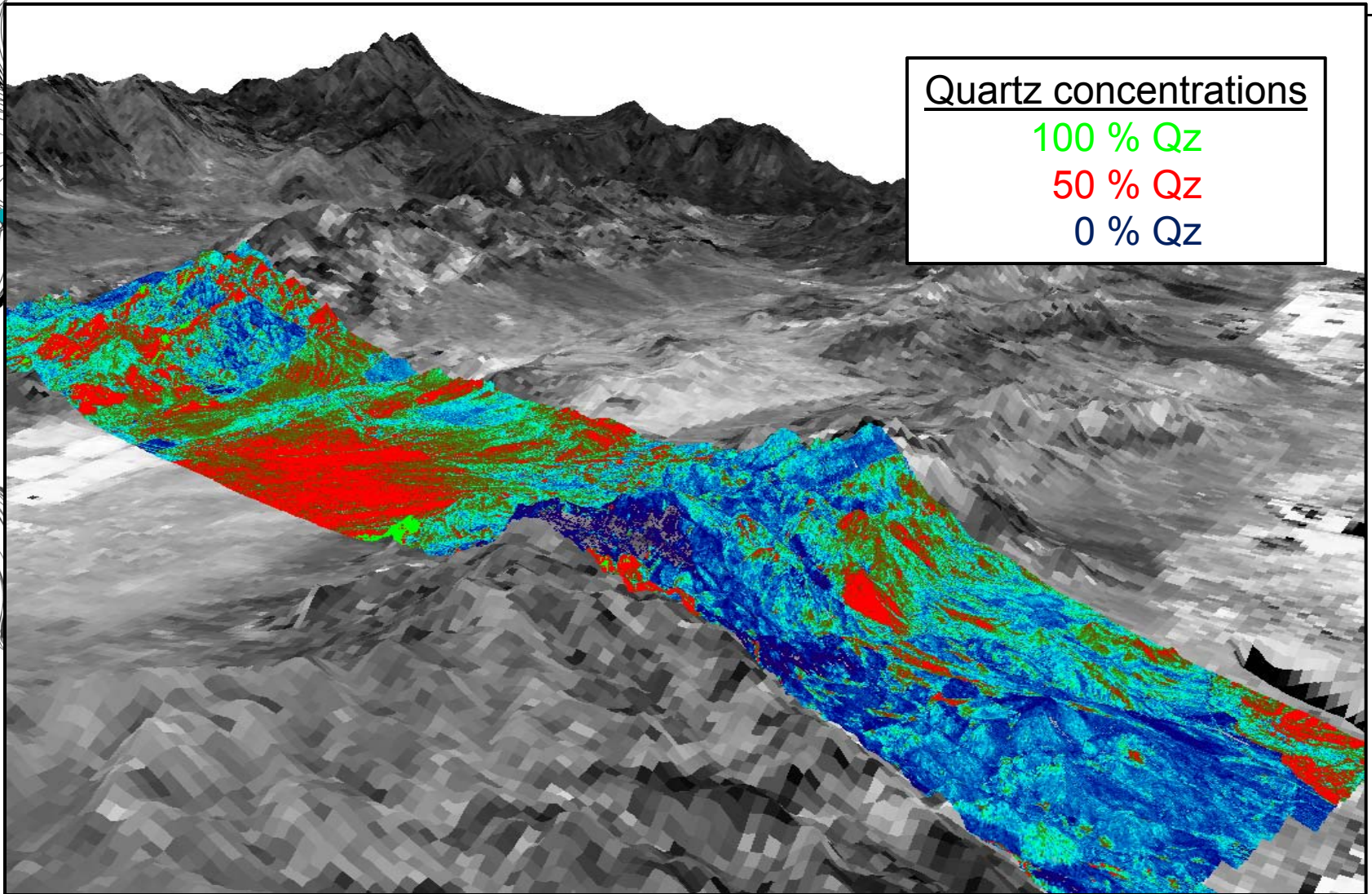
QUANTITATIVE AIRBORNE ANALYSIS

APPLYING PLS MODEL TO AIRBORNE DATA – QTZ CONCENTRATION AS GRAYSCALE IMAGE



QUANTITATIVE AIRBORNE ANALYSIS

APPLYING PLS MODEL TO AIRBORNE DATA - DENSITY SLICED



Quartz concentrations

100 % Qz

50 % Qz

0 % Qz



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SOURCES OF ADDITIONAL INFORMATION: TEXTBOOKS WITH TIR CHAPTERS

- **C. Kuenzer und S. Dech (Eds.)** Thermal Infrared Remote Sensing: Sensors, Method, Applications (2013)
- **Drury** (2001): Image Interpretation in Geology (3rd Edition); Chapter 6
- **Lillesand & Kiefer** (2000): Remote Sensing and Image Interpretation (4th Edition); Chapter 5
- **Sabins** (1997): Remote Sensing – Principles and Interpretation (3rd Edition); Chapter 5
- **Abrams et al** (2001): Imaging Spectrometry in the Thermal Infrared; in vander Meer & de Jong (2001): Imaging Spectrometry; Chapter 10
- **The Remote Sensing Tutorial**
<http://www.fas.org/irp/imint/docs/rst/>
Section 9
- **Gupta** (2003): Remote Sensing Geology (2nd Edition); Chapter 9



SOURCES OF ADDITIONAL INFORMATION: ARTICLES AND CHAPTERS MENTIONED IN TEXT

- **Hecker et al. (2013)** Thermal Infrared Spectroscopy in the Laboratory and Field in Support of Land Surface Remote Sensing. http://dx.doi.org/10.1007/978-94-007-6639-6_3
- **Riley and Hecker (2013)** Mineral Mapping with Airborne Hyperspectral Thermal Infrared Remote Sensing at Cuprite, Nevada, USA, http://dx.doi.org/10.1007/978-94-007-6639-6_24
- **van der Meer et al. (2012)** Multi - and hyperspectral geologic remote sensing : a review. <http://dx.doi.org/10.1016/j.jag.2011.08.002>
- **Hecker et al. (2012)** Thermal infrared spectroscopy and partial least squares regression to determine mineral modes of granitoid rocks. <http://dx.doi.org/10.1029/2011GC004004>
- **Hecker et al. (2011)** Thermal infrared spectrometer for earth science remote sensing applications : instrument modifications and measurement procedures. <http://dx.doi.org/10.3390/s111110981>
- **Hecker et al. (2010)** Thermal infrared spectroscopy on feldspars : successes, limitations and their implications for remote sensing. <http://dx.doi.org/10.1016/j.earscirev.2010.07.005>

QUESTIONS??



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